

# **Recent results from EXO-200**

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# **Double beta decay**







2v mode: a conventional 2nd order process in Standard Model 0ν mode: a hypothetical process can happen only if:  $<m_v> ≠ 0, v = v$  $|\Delta L|=2, |\Delta(B-L)|=2$ 

To reach high measurement sensitivity for  $0\nu$  mode one requires,

- High energy resolution
- · Large Isotope mass
- Low background

Simulated double beta decay spectrum P.Vogel. arXiv:hep-ph/0611243



# Why xenon

Energy resolution is poorer than the crystalline devices (~ factor 10), but...
Monolithic detector. Xenon can form detection medium, allow self shielding, surface contamination minimized. Very good for large scale detectors.
Has high Q value. Located in a region relatively free from natural radioactivity.

- Isotopic enrichment is easier. Xe is already a gas & <sup>136</sup>Xe is the heaviest isotope.
- Xenon is "reusable". Can be purified & recycled into new detector (no crystal growth).
- Minimal cosmogenic activation. No long lived radioactive isotopes of Xe. Energy resolution in LXe can be improved. Scintillation light/ionization correlation.
- Particle identification. Slightly limited, but can be used to tag alphas from Rn chain.

... admits a novel coincidence technique. Background reduction by Ba

daughter tagging.

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## EXO-200 detector



# The EXO-200 TPC



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Two almost identical halves reading ionization and 178 nm scintillation, each with:

- 38 U triplet wire channels (charge)
- 38 V triplet wire channels, crossed at 60° (induction)
- 234 large area avalanche photodiodes (APDs, light in groups of 7)
- Wire pitch 3 mm (9 mm per channel)
- Wire planes 6 mm apart and 6 mm from APD plane
- All signals digitized at 1 MS/s, ±1024S around trigger
- Drift field 376 V/cm
  - Field shaping rings: copper
  - Supports: acrylic
  - Light reflectors/diffusers: Teflon
  - APD support plane: copper; Au (Al) coated for contact (light reflection)
  - Central cathode, U+V wires: photo-etched phosphor bronze
  - Flex cables for bias/readout: copper on kapton, no glue

Comprehensive material screening program

Goal: 40 cnts/2y in  $0\nu\beta\beta \pm 2\sigma$  ROI, 140 kg LXe

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Copper vessel 1.37 mm thick 175 kg LXe, 80.6% enr. in <sup>136</sup>Xe Copper conduits (6) for:

- APD bias and readout cables
- U+V wires bias and readout

 LXe supply and return Epoxy feedthroughs at cold and warm doors Dedicated HV bias line

EXO-200 detector: Materials screening:

JINST 7 (2012) P05010 Characterization of APDs: NIM A608 68-75 (2009) NIM A591, 490-509 (2008)

### Muon veto

ESSINGTON

50 mm thick plastic scintillator panels surrounding TPC on four sides. 95.5 ± 0.6 % efficiency Veto cuts (8.6% combined dead time) CHILLER

- 25 ms after muon veto hit
- 60 s after muon track in TPC
- 1 s after every TPC event

# Data taking phases and xenon purity



Sep 2011 – Hardware upgrades •APD gain increase by factor 2 •improved U-wire shaping •added outer lead shield

#### Purity

Electron lifetime  $\tau_e$  is determined by measuring the attenuation of the ionization signal as a function of drift time for the full-absorption peak of gamma ray sources

For this analysis, the recirculation rate was increased to 14 slpm, leading to long electron lifetimes in the TPC

At  $\tau_e = 3$  ms: •max. drift time ~110 µs •loss of charge is 3.6% at full drift length

Ultraclean pump:

Rev Sci Instrum. 82(10):105114 Xenon purity with mass spectroscopy: NIM A675 (2012) 40-46 Gas purity monitors: NIM A659 (2011) 215-228

## **Event reconstruction**

- Signal finding matched filters applied on U,V and APDs waveforms
- Signal parameter estimation (t, E) for charge and light
- Cluster finding assignment to Single Site (SS) or Multiple Site (MS): resolution 18mm in X and Y and 6 mm in Z

Amplitudes corrected by channel for gain variation Signal fitting functions use individual parameters for each channel Optimized light correction using charge position Charge corrected for inefficiency on small drift Require events to be fully reconstructed in 3D

Reconstruction efficiency for  $0\nu\beta\beta$  is 71% – estimated by MC and verified by comparing the  $2\nu\beta\beta$  MC efficiency with low background data, over a broad range in energy

SS and MS spectra are fitted simultaneously with MC-generated probability density functions

# **Combining ionization and scintillation**



# **Calibrations**



# **Calibrations source spectrum agreement**



•Multi site (MS) and single site (SS) data (black points) are compared to model (blue curves)

- Single site fraction agrees to within 10%
- Can measure source activities better than 4%

# **Energy** Calibration



Using quadratic model for energy calibration, single- and multi-site residual are < 0.1%

Energy resolution model:  $\sigma_{Tot}^2 = p_0^2 E + p_1^2 + p_2^2 E^2$ 

Resolution dominated by constant (noise) term p<sub>1</sub>

At  $Q_{\beta\beta}$  (2458 keV):  $\sigma/E = 1.84 \%$  (SS)  $\sigma/E = 1.93 \%$  (MS)

## Data analysis cuts



- Veto, noise, etc. anticoincidences (5.6% total dead-time)
- «Diagonal» cut to remove alphas and events with bad charge collection
- Only events with 1 scintillation and 3 coordinates reconstructed
- 700 keV low energy cut
- Fiducial cut (15 mm from cathode, 10 mm from anodes and 30 mm from teflon reflector)
- 58% efficiency for 2nu estimated by MC

# Updated results for $2\beta 2\nu$



 $T_{1/2}$  = (2.172 ± 0.017 stat ± 0.06 sys)·10<sup>21</sup> yr The most accurate  $T_{1/2}$  of any 2νββ decay (arXiv:1306.6106)

> KamLAND-Zen  $T_{1/2} = (2.38 \pm 0.02 \text{ stat} \pm 0.14 \text{ sys}) \cdot 10^{21} \text{ yr}$ A.Gando et al., Phys. Rev. C 85 (2012) 045504

# Comparison of the results for $2\beta 2\nu$

Comparison between this result (improved Run 2a) and EXO-200 (2011, Run 1) and KamLAND-Zen (2012) results.

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#### List of the most precise measurements of 2β. Only direct experiments shown.

Nuclide	$T_{1/2}^{2 u\beta\beta} \pm stat \pm sys$	rel. uncert.	$G^{2\nu}$	$M^{2\nu}$	rel. uncert.	Experiment (year)	
	[y]	[%]	$[10^{-21} \text{ y}^{-1}]$	$[{ m MeV}^{-1}]$	[%]		
<sup>136</sup> Xe	$2.172 \pm 0.017 \pm 0.060 \cdot 10^{21}$	$\pm 2.85$	1433	0.0217	±1.4	EXO-200 (	this work)
<sup>76</sup> Ge	$1.84^{+0.09+0.11}_{-0.08-0.06}\cdot10^{21}$	$^{+7.7}_{-5.4}$	48.17	0.129	$^{+3.9}_{-2.8}$	GERDA	(2013)
$^{130}$ Te	$7.0\pm 0.9\pm 1.1\cdot 10^{20}$	$\pm 20.3$	1529	0.0371	$\pm 10.2$	NEMO-3	(2011)
$^{116}Cd$	$2.8\pm 0.1\pm 0.3\cdot 10^{19}$	$\pm 11.3$	2764	0.138	$\pm 5.7$	NEMO-3	(2010)
$^{48}Ca$	$4.4^{+0.5}_{-0.4}\pm0.4\cdot10^{19}$	$^{+14.6}_{-12.9}$	15550	0.0464	+7.3 -6.4	NEMO-3	(2010)
<sup>96</sup> Zr	$2.35 \pm 0.14 \pm 0.16 \cdot 10^{19}$	$\pm 9.1$	6816	0.0959	$\pm 4.5$	NEMO-3	(2010)
<sup>150</sup> Nd	$9.11^{+0.25}_{-0.22}\pm0.63\cdot10^{18}$	$^{+7.4}_{-7.3}$	36430	0.0666	$+3.7 \\ -3.7$	NEMO-3	(2009)
$^{100}Mo$	$7.11 \pm 0.02 \pm 0.54 \cdot 10^{18}$	$\pm 7.6$	3308	0.250	$\pm 3.8$	NEMO-3	(2005)
$^{82}$ Se	$9.6 \pm 0.3 \pm 1.0 \cdot 10^{19}$	$\pm 10.9$	1596	0.0980	$\pm 5.4$	NEMO-3	(2005)

# Results for $2\beta 0v$



Low background run 2a No signal observed

Background in ± 1 ROI: 1.5·10<sup>-3</sup> ± 0.1 kg<sup>-1</sup>yr<sup>-1</sup>keV<sup>-1</sup> - perfectly low

T<sup>0vββ</sup> (<sup>136</sup>Xe) > 1.6·10<sup>25</sup> yr (90% C.L.) Phys. Rev. Lett. V.109 I.3 (2012)

Majorana mass limit <140-380 meV



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# Limits on Ονββ



EXO-200 contradicts Klapdor-Kleingrothaus and Krivosheina claim for  $0\nu\beta\beta$  discovery at the 90% for most matrix elements

A. Gando et al. Phys. Rev. C 85 (2012) 045504 H.V. Klapdor-Kleingrothaus et.al. Eur. Phys. J. A12 (2001) 147

H.V. Klapdor-Kleingrothaus and I.V. Krivosheina, Mod. Phys. Lett., A21 (2006) 1547.

KamLAND-Zen PRL, 110, 062502 (2012)

# Summary

EXO-200 is taking low background data, approved at least end of 2014
About 3 times larger dataset acquired
Detector working well, met our goals:
Energy resolution: 1.84% at Q<sub>ββ</sub>
Background: 1.5 x 10<sup>-3</sup> kg<sup>-1</sup>keV<sup>-1</sup>yr<sup>-1</sup>
1 (5) counts in 1σ (2σ) 0vββ ROI

 $T_{1/2}^{0\nu\beta\beta} > 1.6 \cdot 10^{25} \text{ yr}$  $\langle m_{\beta\beta} \rangle < 140-380 \text{ meV}$ (90% C.L.)

Phys. Rev. Lett. V.109 I.3 (2012)







## nEXO at SNOLab



Conceptual design for SNOLab

# EXO-200 and nEXO projected sensitivities



Blue bands are 68%CL from oscillation experiments for "Inverted" and "Normal" Hierarchy

The EXO-200 "Present limit" is the 90%CL envelope of Limits (for different NMEs) from PRL 109 (2012) 032505

The EXO-200 "Ultimate" sensitivity: 90%CL for no signal in 4 yrs livetime with new analysis & Rn removal

The "Initial nEXO" band refers to a detector directly scaled from EXO-200, including its measured background and 10yr livetime.

The "Final nEXO" band refers to the same detector and no background other than 2v

### The EXO Collaboration





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# Rn Content in Xenon



Using the Bi-Po (Rn daughter) coincidence technique, we can estimate the Rn content in our detector. The <sup>214</sup>Bi decay rate is consistent with measurements from alpha-spectroscopy.

Long-term study shows a constant source of <sup>222</sup>Rn dissolving in <sup>enr</sup>LXe: 360  $\pm$  65  $\mu$ Bq (Fid. vol.)

### EXO-200: the first 200kg Double Beta Decay Experiment



#### Centrifuge facility in Russia





Enriched xenon storage bottles for EXO

EXO collaboration currently have 200 kg of xenon enriched to 80% =160 kg of <sup>136</sup>Xe

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# The EXO-200 TPC

**APDs** 







**Signal cables** 

**Charge detection wires** 

# Background counts in $\pm 1,2 \sigma$ ROI



EXO-200 goal (slide 3):

40 cnts/2y in ±2σ ROI, 140 kg LXe

In this data 120 days, 98.5 kg, this would be: 4.6

	Expected events from fit					
	±1	Lσ	±2	2 σ		
<sup>222</sup> Rn in cryostat air-gap	1.9	±0.2	2.9	±0.3		
<sup>238</sup> U in LXe Vessel	0.9	±0.2	1.3	±0.3		
<sup>232</sup> Th in LXe Vessel	0.9	±0.1	2.9	±0.3		
<sup>214</sup> Bi on Cathode	0.2	±0.01	0.3	±0.02		
All Others	~0.2		~0.2			
Total	4.1	±0.3	7.5	±0.5		
Observed	1		5			
Background index b (kg <sup>-1</sup> yr <sup>-1</sup> keV <sup>-1</sup> )	1.5·10	<sup>-3</sup> ± 0.1	1.4.10	<sup>-3</sup> ± 0.1		

Expected from the fit: 7.5

**Observed: 5** 

Background within expectation

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#### KamLand-ZEN



FIG. 4: Energy spectrum of selected  $\beta\beta$  decay candidates together with the best-fit backgrounds and  $2\nu\beta\beta$  decays, and the 90% C.L. upper limit for  $0\nu\beta\beta$  decays.

 $\begin{array}{l} 112.3 \ days \ 38.6 kg-yr \\ T2\nu 1/2 = 2.30 \pm 0.02 (stat) \pm 0.12 (syst) \ \times 10^{21} \ years \ arXiv:1205.6372 \\ T0\nu 1/2 \ > \ 6.2 \ \times 10^{24} \ years \ (KL-Zen \ 112 days) \\ 23.08.2013 \ \langle \ m\beta\beta \ \rangle \ < 0.26 \ \sim 0.54 \ eV \ @ 90\% \ C.L. \end{array}$ 

### Muon track in EXO-200



#### U and V wires

A track from a cosmic-ray muon in EXO-200. The horizontal axis represents time (uncalibrated for now) while the vertical is the wire position (see sketch). V wires see inductive signals while U wires collects the charge.

The muon in the present event traverses the cathode grid, leaving a long track in one TPC module and a shorter one in the other.

### Barium Tagging R&D Summary





One proposed barium tagging scheme

Method	Summary			
RIS probe	Desorb and resonantly ionize Ba from probe tip, then identify with laser spectroscopy in ion trap			
Hot probe	Release neutral by heating and ionize with hot surface, then identify with laser spectroscopy in ion trap			
Solid Xe probe	Storage and spectroscopy of Ba+ in Xe ice			
Direct tag in liquid	Laser identification of Ba+ in liquid Xe			
Gas Xe extraction	Guide ions from high pressure (10 bar) Xe to low pressure trapping region, then identify with laser spectroscopy			