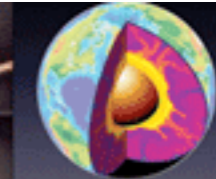
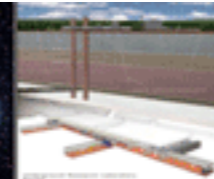


# LAGUNA

Design of a pan-European  
infrastructure for  
Large Apparatus for Grand Unification  
and Neutrino Astrophysics



# LAGUNA/LBNO

**André Rubbia (ETH Zürich)**

**on behalf of LAGUNA-LBNO**

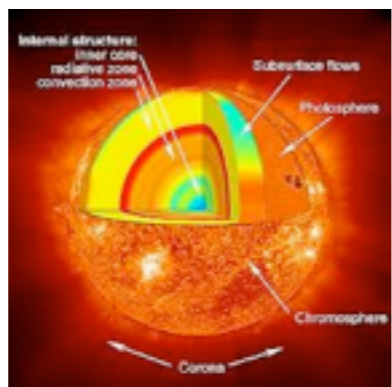
16th LOMONOSOV CONFERENCE ON  
ELEMENTARY PARTICLE PHYSICS



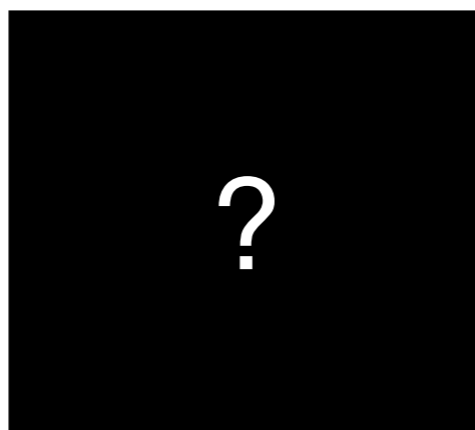
# Goal: neutrino underground observatory



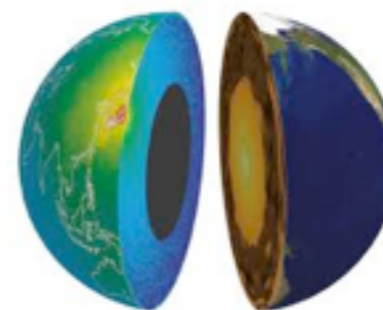
Supernova



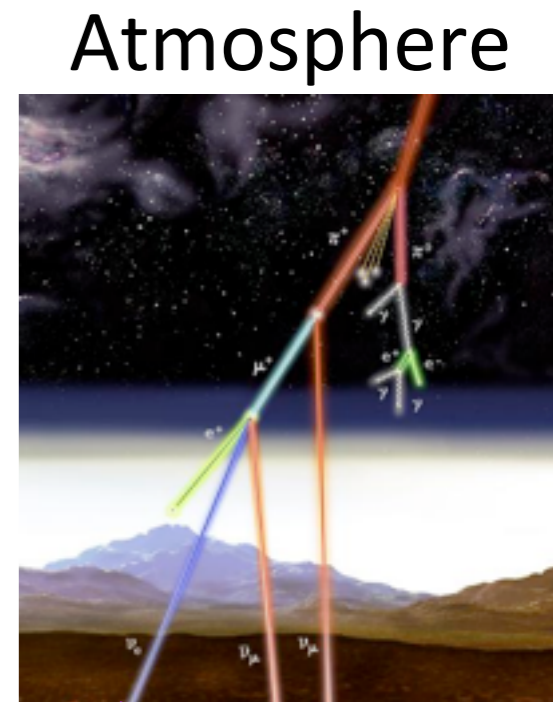
Sun



Unknown ?



Earth



Neutrinos from MeV to 10's GeV

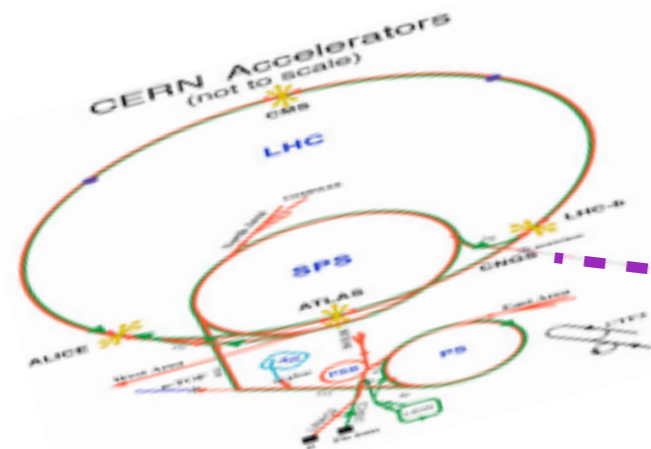
Neutrino oscillations  $\rightarrow$  MH, CPV, precision

Proton lifetime

Address questions of particle and astroparticle physics



Reactors



Accelerators



underground

Deep

Terrestrial baseline

Proton decays

# The LAGUNA history



- **LAGUNA DS** (FP7 Design Study 2008-2011)
  - ~100 members; 10 countries
  - 3 detector technologies  $\otimes$  7 sites, different baselines (130  $\rightarrow$  2300km)
- **LAGUNA-LBNO DS** (FP7 DS Long Baseline Neutrino Oscillations, 2011-2014)
  - ~300 members; 14 countries + CERN
  - Down selection of sites & detectors
- **LENA** (Low Energy Neutrino Astrophysics “Whitepaper”, March 2012)
  - Focused at low energy (MeV) region, with GeV range physics complementary to LBNO
  - ~80 authors; 36 institutions
- **LBNO** (CERN SPSC EoI for a very long baseline neutrino oscillation experiment, June 2012)
  - Consensus towards full long baseline physics + full astroparticle as mandatory physics drivers
  - An incremental approach with clear phase 1 physics capabilities
  - ~230 authors; 51 institutions
  - CERN-SPSC-2012-021 ; SPSC-EOI-007, under review

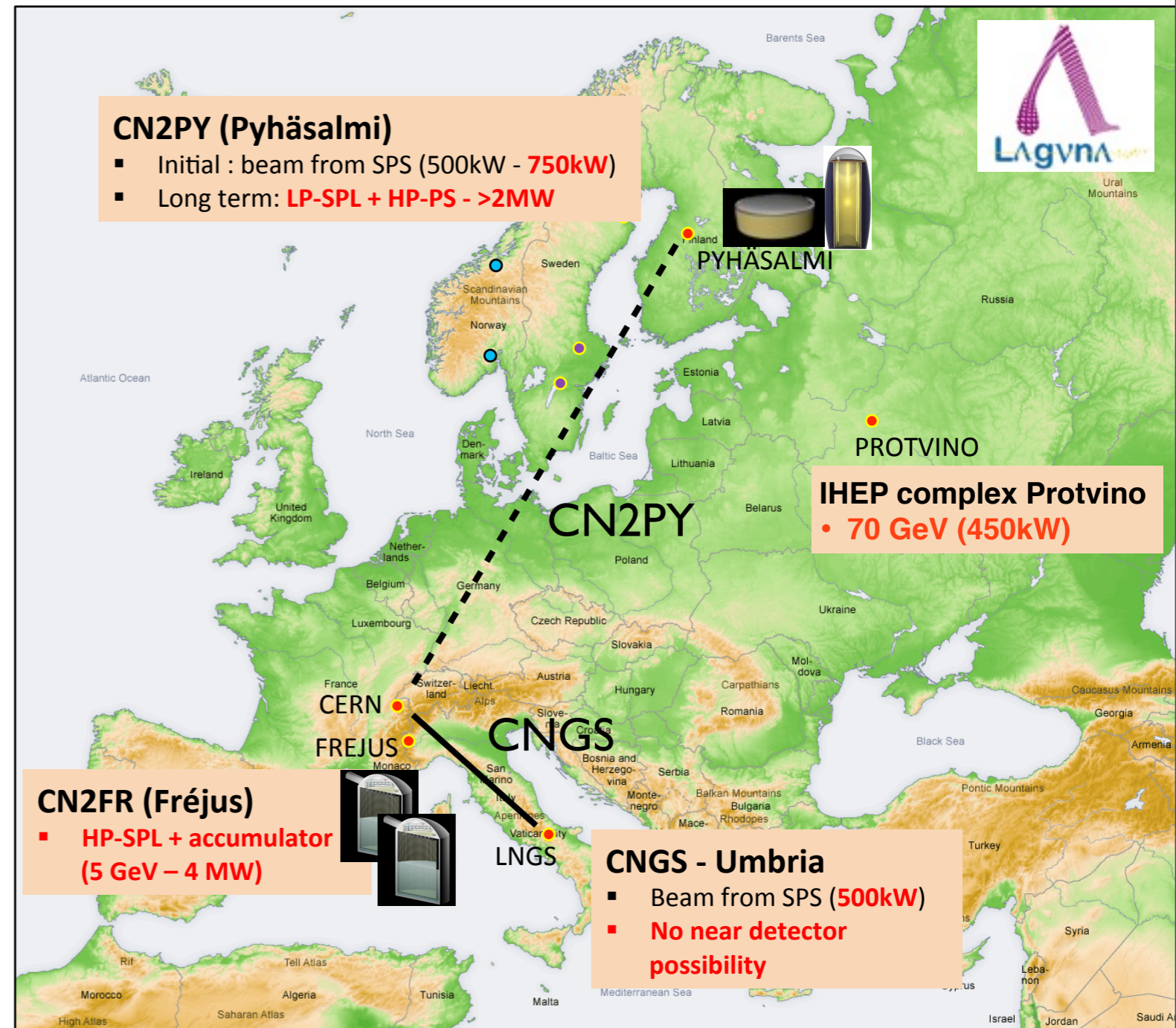
**Large Apparatus for  
Grand Unification  
and Neutrino  
Astrophysics - Long  
Baseline Neutrino  
Oscillations**

# LAGUNA-LBNO: sites overview

## Three far sites considered in details

arXiv:1003.1921 [hep-ph]

- **Option 1: Pyhäsalmi mine** (privately owned), 4000 m.w.e overburden, excellent infrastructure for deep underground access
- **Option 2: Fréjus**, nearby road tunnel, 4800 m.w.e. overburden, horizontal access
- **Option 3: Umbria** (LNGS extension), green site with horizontal access, 2000 m.w.e., CNGS off-axis beam
- **Protons and beams:**
  - Design of new CERN conventional neutrino beam to Finland (CN2PY) Baseline = 2300 km
  - Upgrades of CERN SPS to 700kW
  - New CERN HP-PS (2MW@50 GeV)
  - **Recently:** assessment of a new conventional beam coupled to accelerator upgrade at Protvino, Russia (OMEGA project) – Baseline = 1160 km



- **Detector options:** 20, 50, 100 kton LAr; 50 kton LSc and 540 kton WCD



# LBNO (CERN SPSC-EOI-007)



<http://cdsweb.cern.ch/record/1457543>

- **In June 2012, an enlarged LAGUNA-LBNO Consortium has put forward an Expression of Interest to CERN, focused on neutrino Mass Hierarchy determination and CPV discovery coupled to a full astrophysics programme at the Pyhäsalmi (Finland) site**
  - Based on the findings of several design studies – LAGUNA/LAGUNA-LBNO and EUROv.
  - Supported by rock, civil, detector engineering designs and many years of detector R&D
- **An incremental long-baseline program with a competitive 1st stage guaranteeing high level physics performance from the beginning.**
  - LBNO Stage 1 is based on a 20 kt fid. LAr detector (double phase) and a conventional beam from the CERN SPS of 700 kW at 2300 km.
  - If the findings from Stage 1 require, the detector and the beam will be upgraded to 70 kton mass and 2 MW proton power.
  - The costs, possible implementation schemes and physics potentials will be further studied until the end of 2014.
- **Initial positive feedback from SPSC (108<sup>th</sup> minutes, January 2013)**
  - **The SPSC supports the physics case and recognises its timely relevance in the rapidly evolving neutrino physics landscape.**
  - **SPSC notes that the Finnish Government could not commit to host LAGUNA-LBNO in the proposed Pyhäsalmi site**
  - **SPSC supports double phase LAr TPC as promising technique for future LBL**
  - **SPSC encourages LBNO to proceed with necessary R&D for validation of double phase LAr TPC on large scale**

# LBNO Expression of Interest

A. Stahl,<sup>1</sup> C. Wiebusch,<sup>1</sup> A. M. Guler,<sup>2</sup> M. Kamiscioglu,<sup>2</sup> R. Sever,<sup>2</sup> A.U. Yilmazer,<sup>3</sup> C. Gunes,<sup>3</sup>  
D. Yilmaz,<sup>3</sup> P. Del Amo Sanchez,<sup>4</sup> D. Duchesneau,<sup>4</sup> H. Pessard,<sup>4</sup> E. Marcoulaki,<sup>5</sup> I. A.  
Papazoglou,<sup>5</sup> V. Berardi,<sup>6</sup> F. Cafagna,<sup>6</sup> M.G. Catanesi,<sup>6</sup> L. Magaletti,<sup>6</sup> A. Mercadante,<sup>6</sup>  
M. Quinto,<sup>6</sup> E. Radicioni,<sup>6</sup> A. Ereditato,<sup>7</sup> I. Kreslo,<sup>7</sup> C. Pistillo,<sup>7</sup> M. Weber,<sup>7</sup> A. Ariga,<sup>7</sup> T. Ariga,<sup>7</sup>  
T. Strauss,<sup>7</sup> M. Hierholzer,<sup>7</sup> J. Kawada,<sup>7</sup> C. Hsu,<sup>7</sup> S. Haug,<sup>7</sup> A. Jipa,<sup>8</sup> I. Lazanu,<sup>8</sup> A. Cardini,<sup>9</sup>  
A. Lai,<sup>9</sup> R. Oldeman,<sup>10</sup> M. Thomson,<sup>11</sup> A. Blake,<sup>11</sup> M. Prest,<sup>12</sup> A. Auld,<sup>13</sup> J. Elliot,<sup>13</sup> J. Lumbard,<sup>13</sup>  
C. Thompson,<sup>13</sup> Y.A. Gornushkin,<sup>14</sup> S. Pascoli,<sup>15</sup> R. Collins,<sup>16</sup> M. Haworth,<sup>16</sup> J. Thompson,<sup>16</sup>  
G. Bencivenni,<sup>17</sup> D. Domenici,<sup>17</sup> A. Longhin,<sup>17</sup> A. Blondel,<sup>18</sup> A. Bravar,<sup>18</sup> F. Dufour,<sup>18</sup> Y. Karadzhov,<sup>18</sup>  
A. Korzenev,<sup>18</sup> E. Noah,<sup>18</sup> M. Ravonel,<sup>18</sup> M. Rayner,<sup>18</sup> R. Asfandiyarov,<sup>18</sup> A. Haesler,<sup>18</sup>  
C. Martin,<sup>18</sup> E. Scantamburlo,<sup>18</sup> F. Cadoux,<sup>18</sup> R. Bayes,<sup>19</sup> F.J.P. Soler,<sup>19</sup> L. Aalto-Setälä,<sup>20</sup>  
K. Enqvist,<sup>20</sup> K. Huitu,<sup>20</sup> K. Rummukainen,<sup>20</sup> G. Nuijten,<sup>21</sup> K.J. Eskola,<sup>22</sup> K. Kainulainen,<sup>22</sup>  
T. Kalliokoski,<sup>22</sup> J. Kumpulainen,<sup>22</sup> K. Loo,<sup>22</sup> J. Maalampi,<sup>22</sup> M. Manninen,<sup>22</sup> I. Moore,<sup>22</sup>  
J. Suhonen,<sup>22</sup> W.H. Trzaska,<sup>22</sup> K. Tuominen,<sup>22</sup> A. Virtanen,<sup>22</sup> I. Bertram,<sup>23</sup> A. Finch,<sup>23</sup> N. Grant,<sup>23</sup>  
L.L. Kormos,<sup>23</sup> P. Ratoff,<sup>23</sup> G. Christodoulou,<sup>24</sup> J. Coleman,<sup>24</sup> C. Touramanis,<sup>24</sup> K. Mavrokoridis,<sup>24</sup>  
M. Murdoch,<sup>24</sup> N. McCauley,<sup>24</sup> D. Payne,<sup>24</sup> P. Jonsson,<sup>25</sup> A. Kaboth,<sup>25</sup> K. Long,<sup>25</sup> M. Malek,<sup>25</sup>  
M. Scott,<sup>25</sup> Y. Uchida,<sup>25</sup> M.O. Wascko,<sup>25</sup> F. Di Lodovico,<sup>26</sup> J.R. Wilson,<sup>26</sup> B. Still,<sup>26</sup> R. Sacco,<sup>26</sup>  
R. Terri,<sup>26</sup> M. Campanelli,<sup>27</sup> R. Nichol,<sup>27</sup> J. Thomas,<sup>27</sup> A. Izmaylov,<sup>28</sup> M. Khabibullin,<sup>28</sup>  
A. Khotjantsev,<sup>28</sup> Y. Kudenko,<sup>28</sup> V. Matveev,<sup>28</sup> O. Mineev,<sup>28</sup> N. Yershov,<sup>28</sup> V. Palladino,<sup>29</sup> J. Evans,<sup>30</sup>  
S. Söldner-Rembold,<sup>30</sup> U.K. Yang,<sup>30</sup> M. Bonesini,<sup>31</sup> T. Pihlajaniemi,<sup>32</sup> M. Weckström,<sup>32</sup> K.  
Mursula,<sup>32</sup> T. Enqvist,<sup>32</sup> P. Kuusiniemi,<sup>32</sup> T. Rähjä,<sup>32</sup> J. Sarkamo,<sup>32</sup> M. Slupecki,<sup>32</sup> J. Hissa,<sup>32</sup> E.  
Kokko,<sup>32</sup> M. Aittola,<sup>32</sup> G. Barr,<sup>33</sup> M.D. Haigh,<sup>33</sup> J. de Jong,<sup>33</sup> H. O’Keeffe,<sup>33</sup> A. Vacheret,<sup>33</sup>  
A. Weber,<sup>33,34</sup> G. Galvanin,<sup>35</sup> M. Temussi,<sup>35</sup> O. Caretta,<sup>34</sup> T. Davenne,<sup>34</sup> C. Densham,<sup>34</sup> J. Ilic,<sup>34</sup>  
P. Loveridge,<sup>34</sup> J. Odell,<sup>34</sup> D. Wark,<sup>34</sup> A. Robert,<sup>36</sup> B. Andrieu,<sup>36</sup> B. Popov,<sup>36,14</sup> C. Giganti,<sup>36</sup>  
J.-M. Levy,<sup>36</sup> J. Dumarchez,<sup>36</sup> M. Buizza-Avanzini,<sup>37</sup> A. Cabrera,<sup>37</sup> J. Dawson,<sup>37</sup> D. Franco,<sup>37</sup>  
D. Kryn,<sup>37</sup> M. Obolensky,<sup>37</sup> T. Patzak,<sup>37</sup> A. Tonazzo,<sup>37</sup> F. Vanucci,<sup>37</sup> D. Orestano,<sup>38</sup> B. Di Micco,<sup>38</sup>  
L. Tortora,<sup>39</sup> O. Bésida,<sup>40</sup> A. Delbart,<sup>40</sup> S. Emery,<sup>40</sup> V. Galymov,<sup>40</sup> E. Mazzucato,<sup>40</sup> G. Vasseur,<sup>40</sup>  
M. Zito,<sup>40</sup> V.A. Kudryavtsev,<sup>41</sup> L.F. Thompson,<sup>41</sup> R. Tsenov,<sup>42</sup> D. Kolev,<sup>42</sup> I. Rusinov,<sup>42</sup>  
M. Bogomilov,<sup>42</sup> G. Vankova,<sup>42</sup> R. Matev,<sup>42</sup> A. Vorobyev,<sup>43</sup> Yu. Novikov,<sup>43</sup> S. Kosyanenko,<sup>43</sup>  
V. Suvorov,<sup>43</sup> G. Gavrilov,<sup>43</sup> E. Baussan,<sup>44</sup> M. Dracos,<sup>44</sup> C. Jollet,<sup>44</sup> A. Meregaglia,<sup>44</sup> E. Vallazza,<sup>45</sup>  
S.K. Agarwalla,<sup>46</sup> T. Li,<sup>46</sup> D. Autiero,<sup>47</sup> L. Chaussard,<sup>47</sup> Y. Déclais,<sup>47</sup> J. Marteau,<sup>47</sup> E. Pennacchio,<sup>47</sup>  
E. Rondio,<sup>48</sup> J. Lagoda,<sup>48</sup> J. Zalipska,<sup>48</sup> P. Przewlocki,<sup>48</sup> K. Grzelak,<sup>49</sup> G. J. Barker,<sup>50</sup> S. Boyd,<sup>50</sup>  
P.F. Harrison,<sup>50</sup> R.P. Litchfield,<sup>50</sup> Y. Ramachers,<sup>50</sup> A. Badertscher,<sup>51</sup> A. Curioni,<sup>51</sup> U. Degunda,<sup>51</sup>  
L. Epprecht,<sup>51</sup> A. Gendotti,<sup>51</sup> L. Knecht,<sup>51</sup> S. DiLuise,<sup>51</sup> S. Horikawa,<sup>51</sup> D. Lussi,<sup>51</sup> S. Murphy,<sup>51</sup>  
G. Natterer,<sup>51</sup> F. Petrollo,<sup>51</sup> L. Periale,<sup>51</sup> A. Rubbia,<sup>51,\*</sup> F. Sergiampietri,<sup>51</sup> and T. Viant<sup>51</sup>

1. III. Physikalisches Institut, RWTH Aachen, Aachen, [Germany](#)
2. Middle East Technical University (METU), Ankara, [Turkey](#)
3. Ankara University, Ankara, [Turkey](#)
4. LAPP, Université de Savoie, CNRS/IN2P3, F-74941 Annecy-le-Vieux, [France](#)
5. Institute of Nuclear Technology-Radiation Protection, National Centre for Scientific Research "Demokritos", Athens, [Greece](#)
6. INFN and Dipartimento interateneo di Fisica di Bari, Bari, [Italy](#)
7. University of Bern, Albert Einstein Center for Fundamental Physics, Laboratory for High Energy Physics (LHEP), Bern, [Switzerland](#)
8. Faculty of Physics, University of Bucharest, Bucharest, [Romania](#)
9. INFN Sezione di Cagliari, Cagliari, [Italy](#)
10. INFN Sezione di Cagliari and Università di Cagliari, Cagliari, [Italy](#)
11. University of Cambridge, Cambridge, [United Kingdom](#)
12. Università dell’Insubria, sede di Como/ INFN Milano Bicocca, Como, [Italy](#)
13. Alan Auld Engineering, Doncaster, [United Kingdom](#)
14. Joint Institute for Nuclear Research, Dubna, Moscow Region, [Russia](#)
15. Institute for Particle Physics Phenomenology, Durham University, [United Kingdom](#)
16. Technodyne International Limited, Eastleigh, Hampshire, [United Kingdom](#)
17. INFN Laboratori Nazionali di Frascati, Frascati, [Italy](#)
18. University of Geneva, Section de Physique, DPNC, Geneva, [Switzerland](#)
19. University of Glasgow, Glasgow, [United Kingdom](#)
20. University of Helsinki, Helsinki, [Finland](#)
21. Rockplan Ltd., Helsinki, [Finland](#)
22. Department of Physics, University of Jyväskylä, [Finland](#)
23. Physics Department, Lancaster University, Lancaster, [United Kingdom](#)
24. University of Liverpool, Department of Physics, Liverpool, [United Kingdom](#)
25. Imperial College, London, [United Kingdom](#)
26. Queen Mary University of London, School of Physics, London, [United Kingdom](#)
27. Dept. of Physics and Astronomy, University College London, London, [United Kingdom](#)
28. Institute for Nuclear Research of the Russian Academy of Sciences, Moscow, [Russia](#)
29. INFN Sezione di Napoli and Università di Napoli, Dipartimento di Fisica, Napoli, [Italy](#)
30. University of Manchester, Manchester, [United Kingdom](#)
31. INFN Milano Bicocca, Milano, [Italy](#)
32. University of Oulu, Oulu, [Finland](#)
33. Oxford University, Department of Physics, Oxford, [United Kingdom](#)
34. STFC, Rutherford Appleton Laboratory, Harwell Oxford, [United Kingdom](#)
35. AGT Ingegneria S.r.l., Perugia, [Italy](#)
36. UPMC, Université Paris Diderot, CNRS/IN2P3, Laboratoire de Physique Nucléaire et de Hautes Energies (LPNHE), Paris, [France](#)
37. APC, AstroParticule et Cosmologie, Université Paris Diderot, CNRS/IN2P3, CEA/Irfu, Observatoire de Paris, Sorbonne Paris Cité Paris, [France](#)
38. Università and INFN Roma Tre, Roma, [Italy](#)
39. INFN Roma Tre, Roma, [Italy](#)
40. IRFU, CEA Saclay, Gif-sur-Yvette, [France](#)
41. University of Sheffield, Department of Physics and Astronomy, Sheffield, [United Kingdom](#)
42. Department of Atomic Physics, Faculty of Physics, St.Kliment Ohridski University of Sofia, Sofia, [Bulgaria](#)
43. Petersburg Nuclear Physics Institute (PNPI), St-Petersburg, [Russia](#)
44. IPHC, Université de Strasbourg, CNRS/IN2P3, Strasbourg, [France](#)
45. INFN Trieste, Trieste, [Italy](#)
46. IFIC (CSIC & University of Valencia), Valencia, [Spain](#)
47. Université de Lyon, Université Claude Bernard Lyon 1, IPN Lyon (IN2P3), Villeurbanne, [France](#)
48. National Centre for Nuclear Research (NCBJ), Warsaw, [Poland](#)
49. Institute of Experimental Physics, Warsaw University (IFD UW), Warsaw, [Poland](#)
50. University of Warwick, Department of Physics, Coventry, [United Kingdom](#)
51. ETH Zurich, Institute for Particle Physics, Zurich, [Switzerland](#)

# LBNO main physics goals

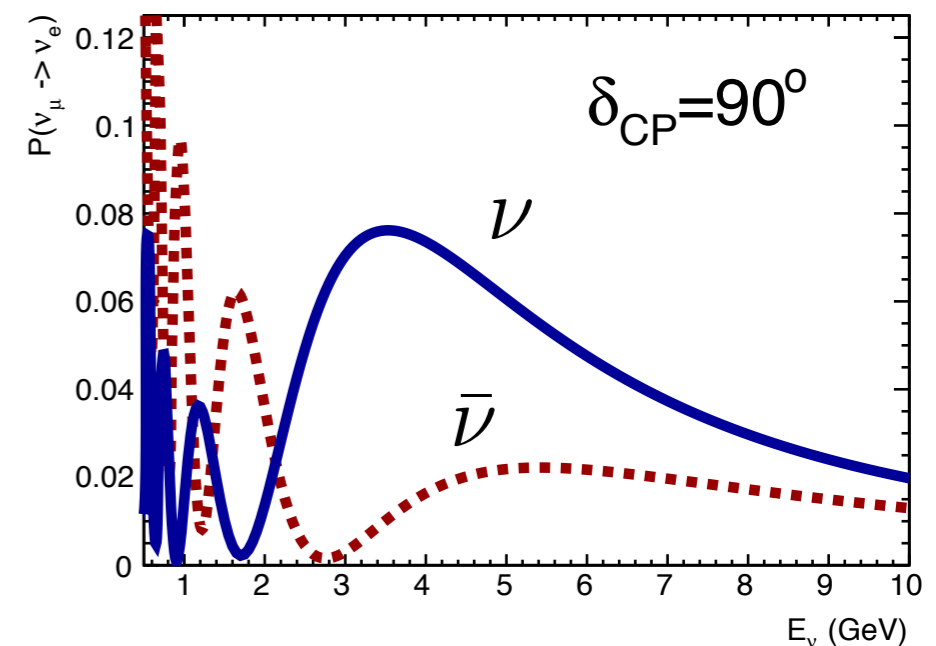
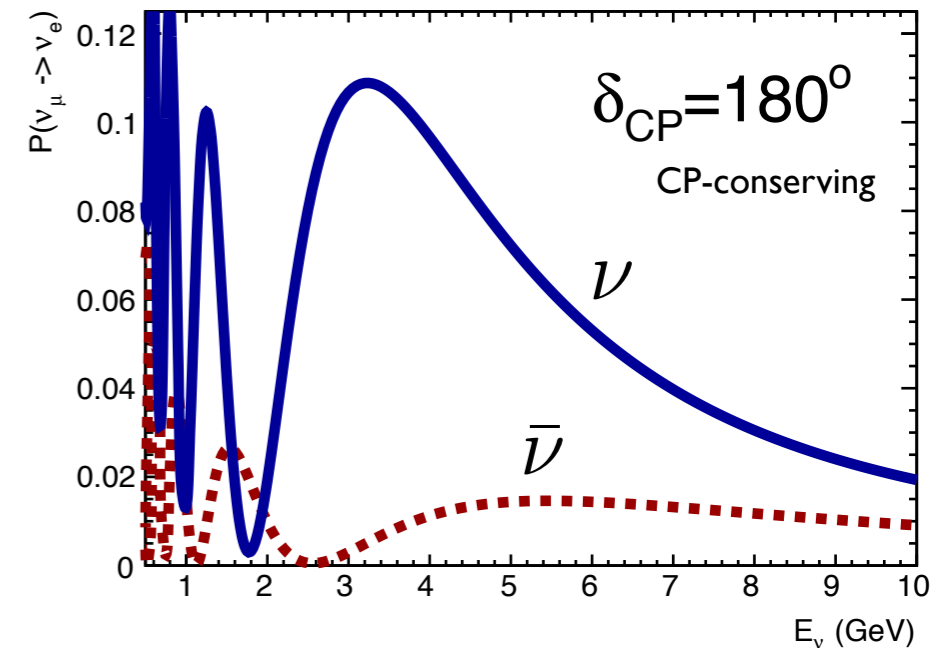
NH

- **Long baseline neutrino oscillations**

- $\nu\mu \rightarrow \nu e$  &  $\nu\mu \rightarrow \nu\tau$  &  $\nu\mu \rightarrow \nu\mu$  &  $\nu\text{NC}$
- Direct measurement of the energy dependence (L/E behaviour) induced by matter effects and CP-phase terms, independently for  $\nu$  and anti- $\nu$ , by direct measurement of event spectrum, in particular covering 1st and 2nd oscillation maxima
- Mass hierarchy determination at  $>5\sigma$  C.L. in first two years of running
- CP-phase measurement and CPV “discovery” ( $\Rightarrow 5\sigma$  C.L.)
- Test of three generation mixing paradigm

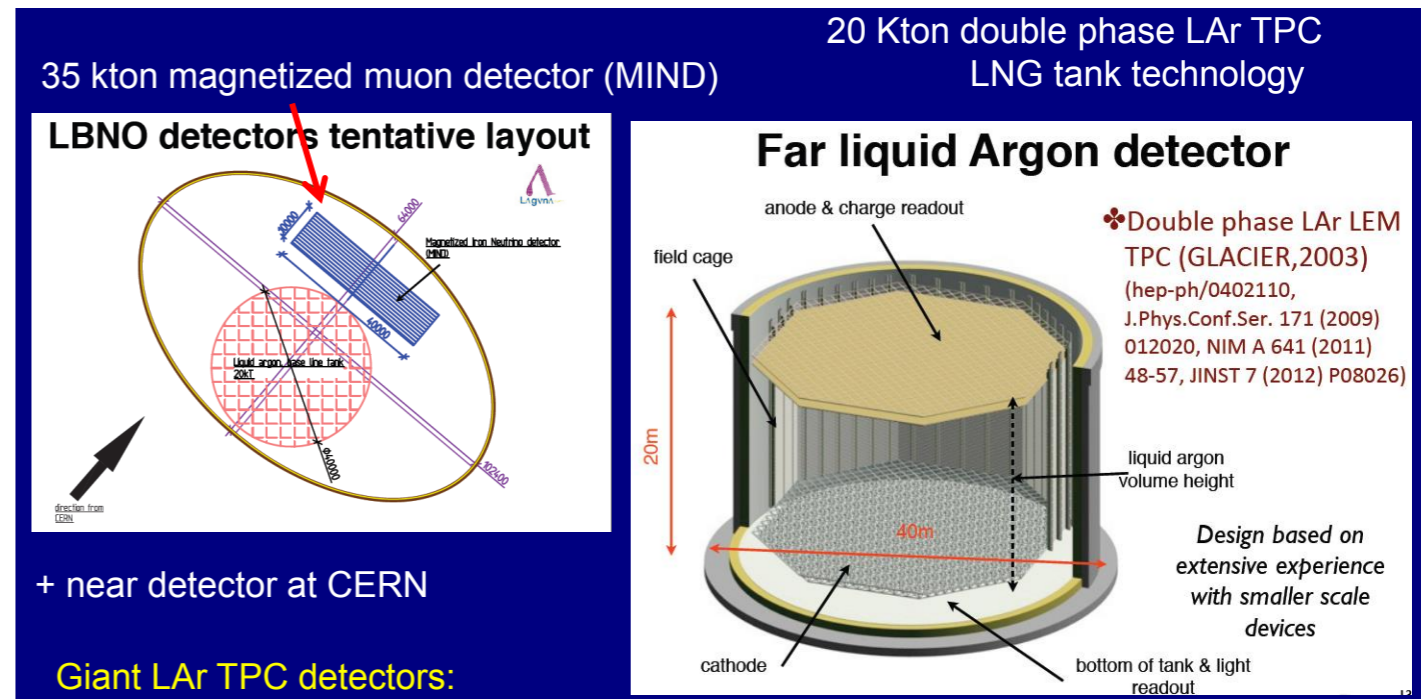
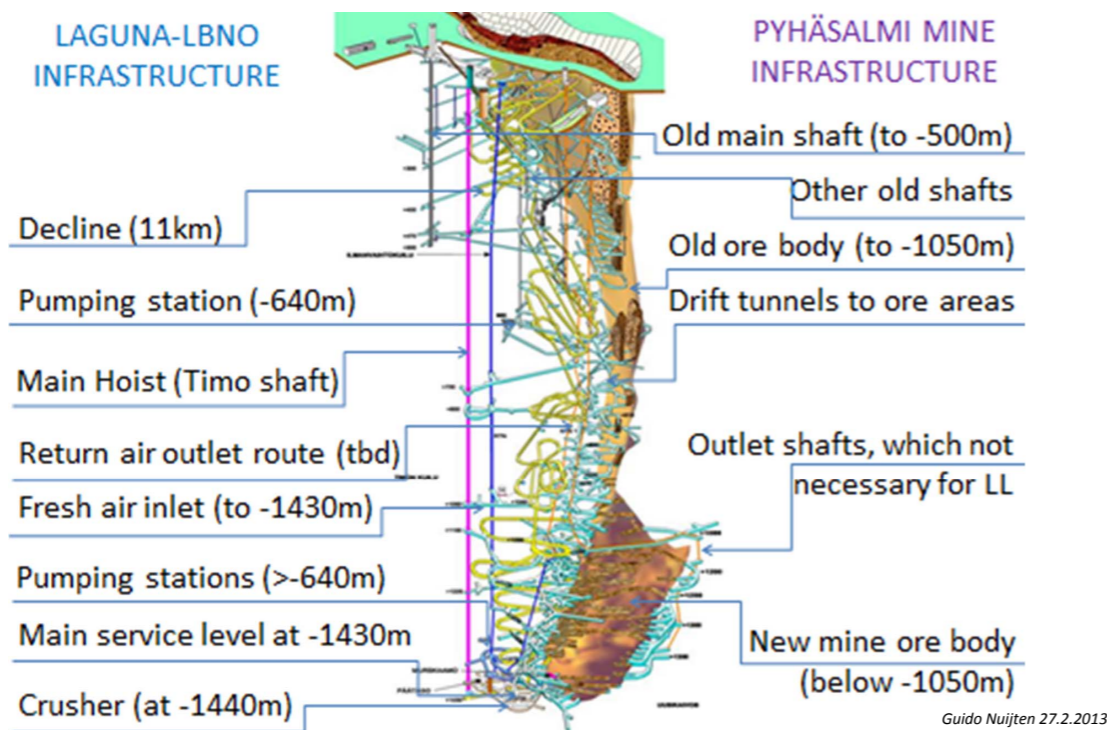
- **A full astrophysics programme**

- Nucleon decays (direct GUT evidence)
- Atmospheric neutrino detection with complementary oscillation measurements and Earth spectroscopy
- Astrophysical neutrino detection and searches for new sources of neutrinos

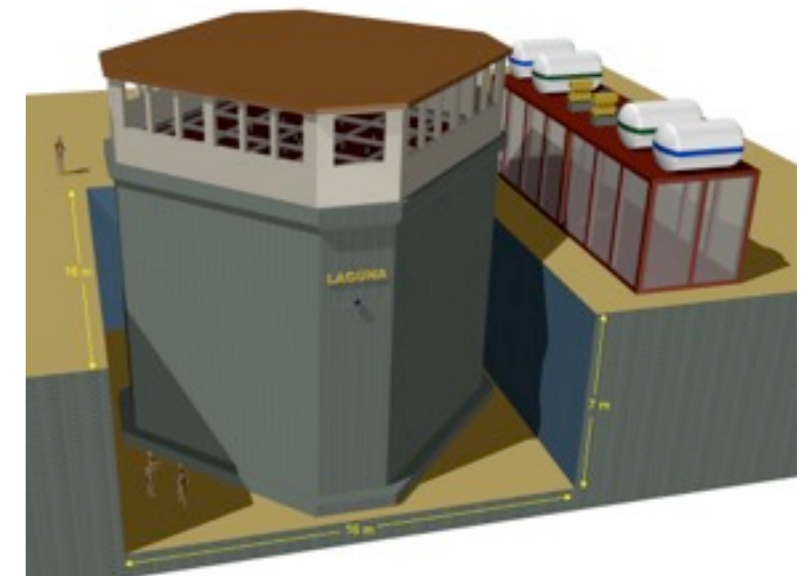


# LBNO far site and detectors

- The LAGUNA LBNO collaboration is in the most advanced state for what concerns all technical implementation and site studies, costing and prototyping.
- The Pyhäsalmi site is extremely convenient (baseline, infrastructures, depth, excavation aspects). An extended site investigation is progressing well (750 m drilled) → Discussions will continue with Finland in order to define its real contribution, after some initial misunderstanding. Alternative sites in Scandinavia are been looked into.



- **Next milestone:** Large-scale LBNO detectors prototyping at CERN, with priority emphasis on a large double-phase LAr demonstrator, using charged-particle test beams (2014-2017).

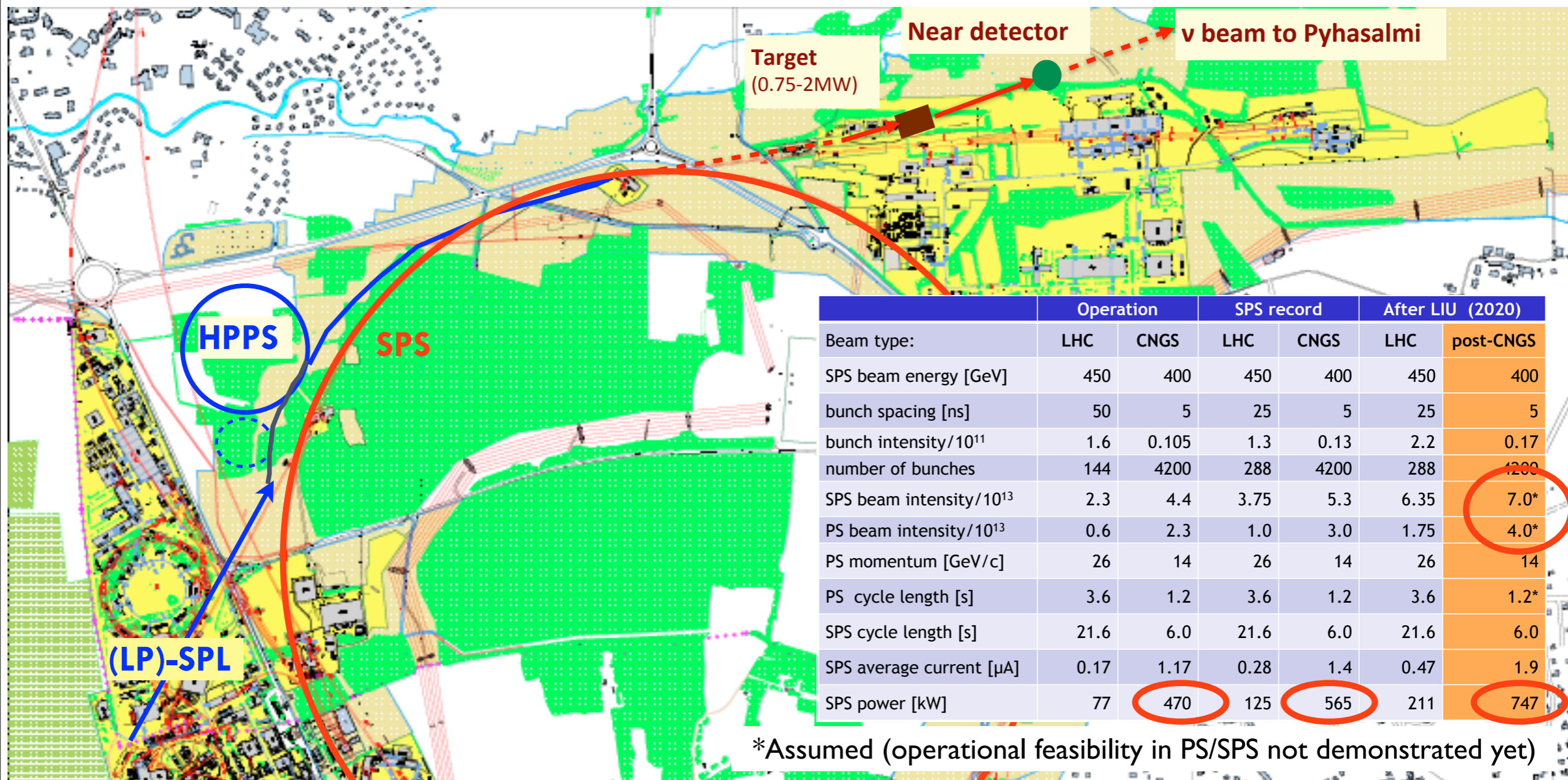




# The CN2PY beam

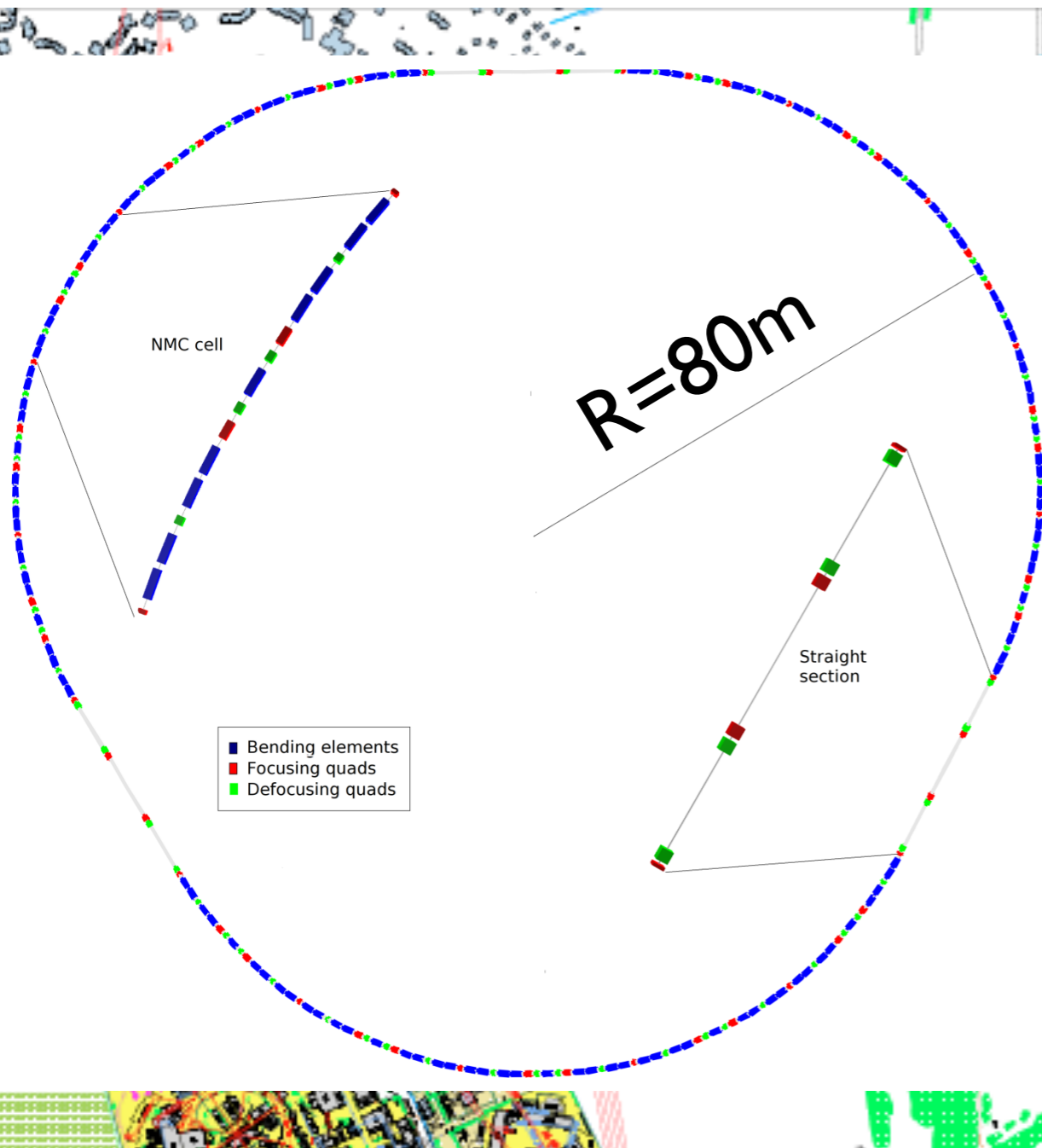


- ▶ **Phase 1** : use the proton beam extracted beam from SPS
  - **400 GeV**, max  $7.0 \cdot 10^{13}$  protons every 6 sec, **750 kW** nominal beam power, 10  $\mu$ s pulse
  - Yearly integrated pot =  $(8-13)e19$  pot / yr depending on “sharing” with other fixed target programmes.
- ▶ **Phase 2** : use the proton beam from the new HP-PS
  - **50(70) GeV**, 1 Hz,  $2.5e14$  ppp, **2 MW** nominal beam power, 4  $\mu$ s pulse

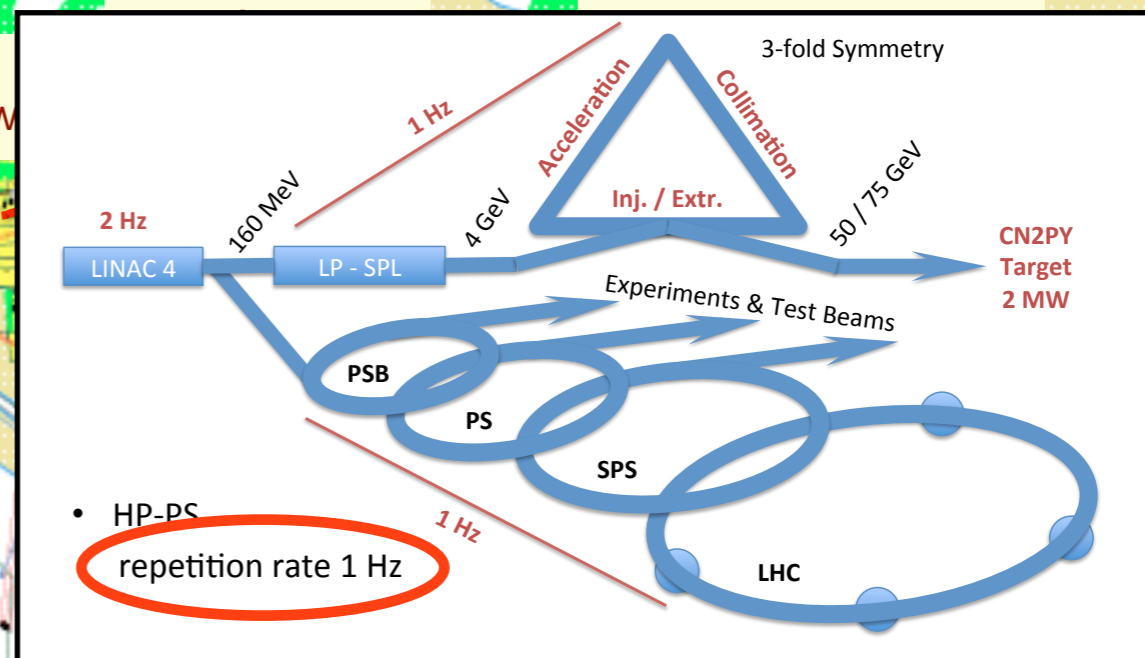


\*Assumed (operational feasibility in PS/SPS not demonstrated yet)

# High power HP-PS study



Target  
(0.75-2MW)

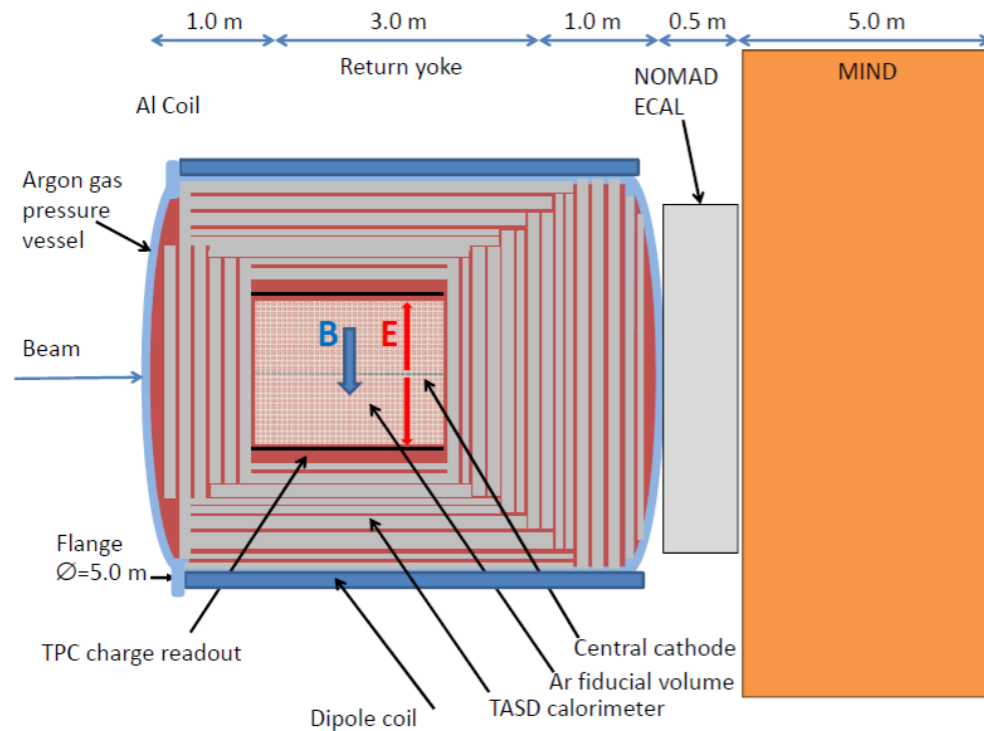


Parameter	50 GeV	75 GeV	Units
Inj. / Extr. Kinetic Energy	4 / 50	4 / 75	[GeV]
Beam power	2		[MW]
Repetition rate	1		[Hz]
$f_{rev} / f_{RF}$ @ inj.	0.248 / 38.97		[MHz]
RF harmonic	157		-
$f_{rev} / f_{RF}$ @ extr.	0.255 / 40.08	0.255 / 40.09	[MHz]
Bunch spacing @ extr.	25		[ns]
Total beam intensity	$2.5 \times 10^{14}$	$1.7 \times 10^{14}$	-
Number of bunches	147		-
Intensity per bunch	$1.7 \times 10^{12}$	$1.25 \times 10^{12}$	-
Main dipole field inj. / extr.	0.17 / 2.1	0.17 / 3.13	[T]
Ramp time	500	500	[ms]
Dipole field rate dB/dt (acc. ramp)	3.9	5.9	[T/s]

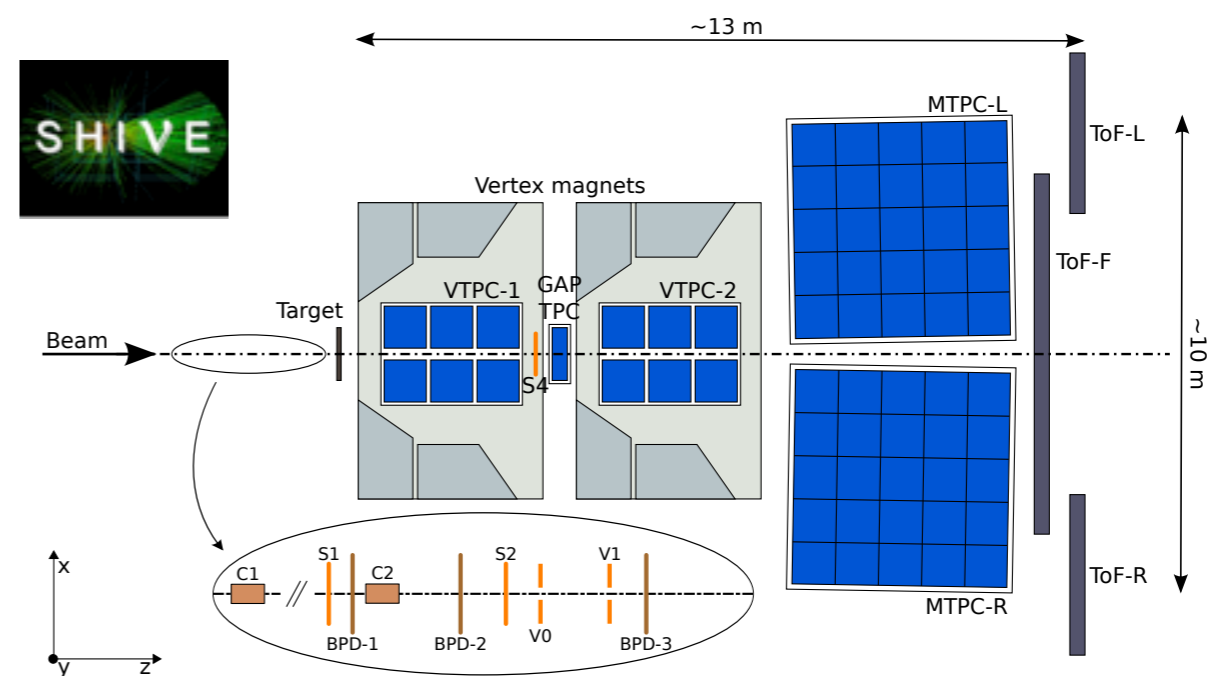
- Basic design well underway and main parameters available
- Optics design well advanced
- Injection and extraction concepts are available
- Basic ideas about accelerating RF system
- Basic ideas about collimation
- Consolidate optics and establish set of requirements for different magnet families.
- Design of magnet foreseen.

# LBNO near detector and hadroproduction

- **Aim:** systematic errors for signal and backgrounds in the far detectors below  $\pm 5\%$ , possibly at the level of  $\pm 2\%$   $\Rightarrow$  control of fluxes, cross-sections, efficiencies,...



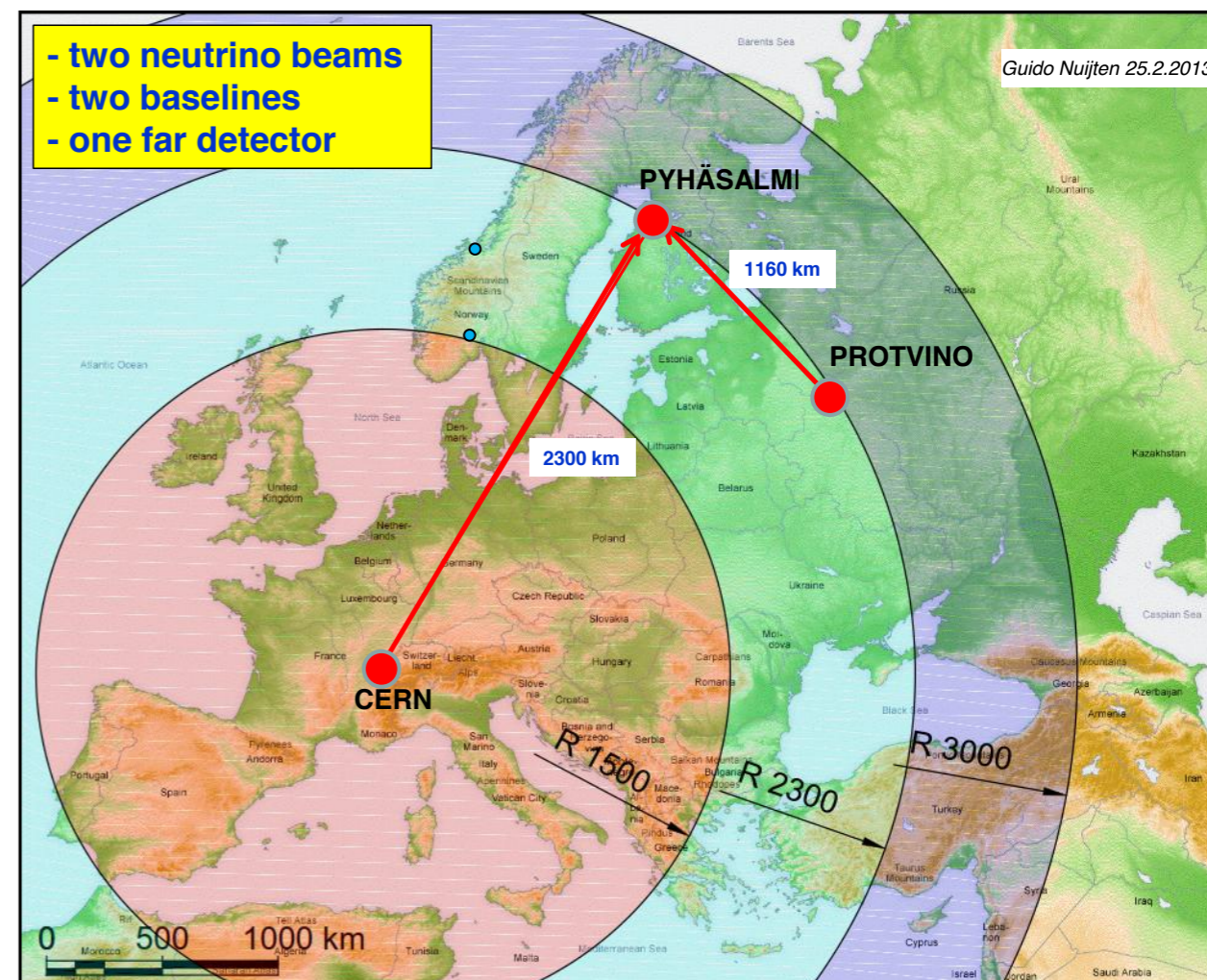
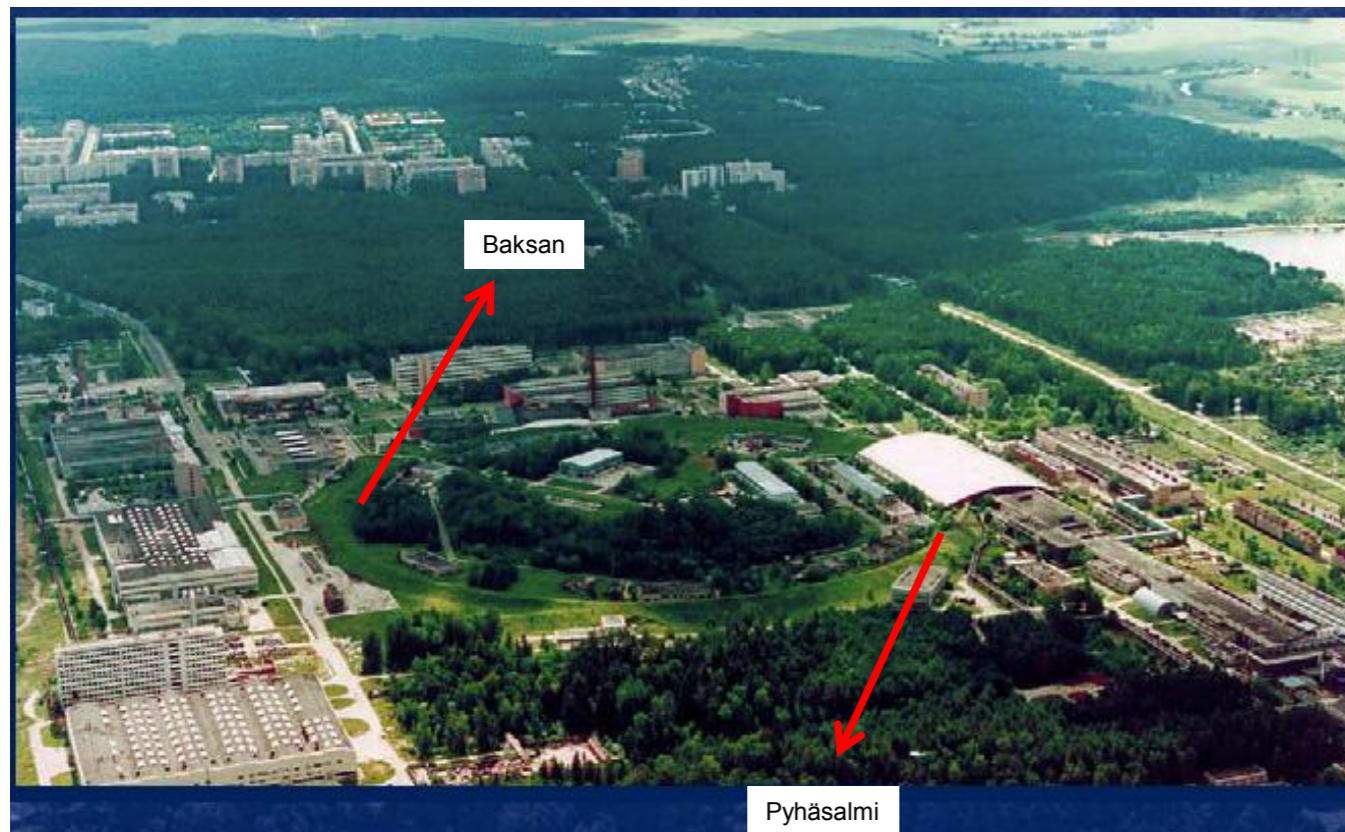
- Concept: 20 bar gas argon-mixture TPC (2.4 m  $\times$  2.4 m  $\times$  3 m) surrounded by scintillator bar tracker embedded in an instrumented magnet with field 0.5T
- 600 kg argon mass in TPC
- 0.2 event/spill @  $7e13$  ppp 400 GeV
- $O(100'000)$  events/year



- It is widely recognized that hadro-production measurements with thin or replica target are really crucial for precision neutrino experiments (eg. K2K, T2K, MINOS).
- CERN NA61 upgrade needed for 400 GeV incident protons

- Precision neutrino cross-section measurements: e.g. MINERVA, T2K-ND280, also nuSTORM

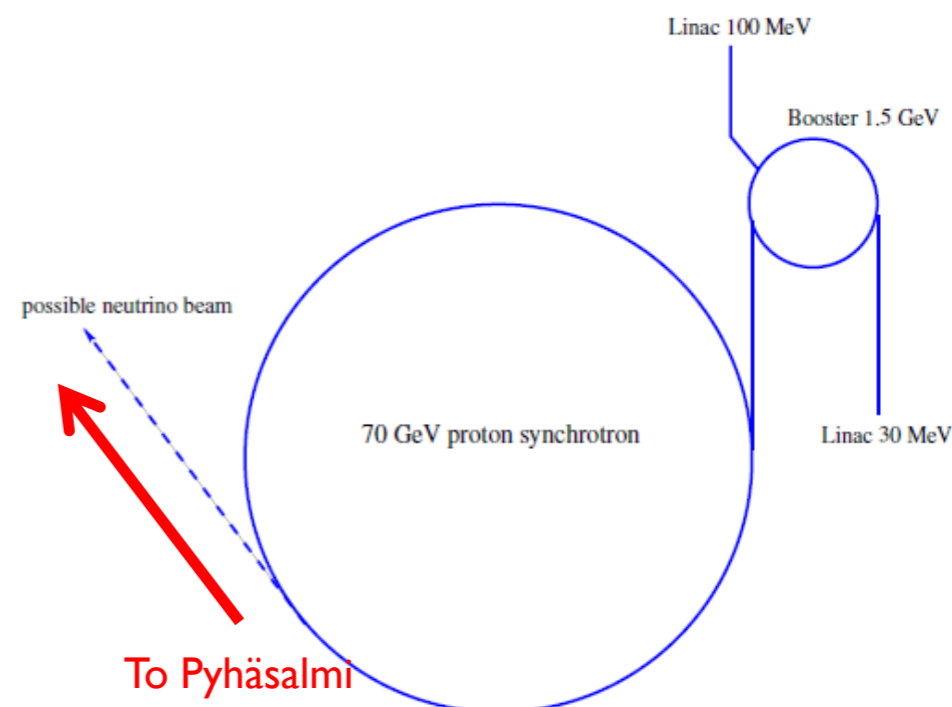
# Possibility of neutrinos from Protvino



## Assumed parameters for neutrino beam:

Proton energy	70 GeV
Repetition rate	0.2 Hz
Intensity	$2.2 \times 10^{14}$ ppp
Power	450 kW
Neutrino channel	200-300 m
Angle to Pyhäsalmi	5.2 deg
Distance to ND	500 - 750 m
ND depth (at 500m)	46 m

**$\approx 1800 \nu_{\mu}$  CC / 20 kton / year (no osc.)**



# Rich MH & CP phenomenology

- First order approximation in expansion (Sato et al.):

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e; L) \simeq & 4c_{13}^2 s_{13}^2 s_{23}^2 \left\{ 1 + \frac{a}{\delta m_{31}^2} \cdot 2(1 - 2s_{13}^2) \right\} \sin^2 \frac{\delta m_{31}^2 L}{4E} \\
 & + c_{13}^2 s_{13} s_{23} \left\{ -\frac{aL}{E} s_{13} s_{23} (1 - 2s_{13}^2) + \frac{\delta m_{21}^2 L}{E} s_{12} (-s_{13} s_{23} s_{12} + c_\delta c_{23} c_{12}) \right\} \sin \frac{\delta m_{31}^2 L}{2E} \\
 & - 4 \frac{\delta m_{21}^2 L}{2E} s_\delta c_{13}^2 s_{13} c_{23} s_{23} c_{12} s_{12} \sin^2 \frac{\delta m_{31}^2 L}{4E}
 \end{aligned}$$

**Matter terms**  $\sim a$   
**CP-even**  
**CP-odd**  $\sim \sin \delta$   
**L/E dependence**

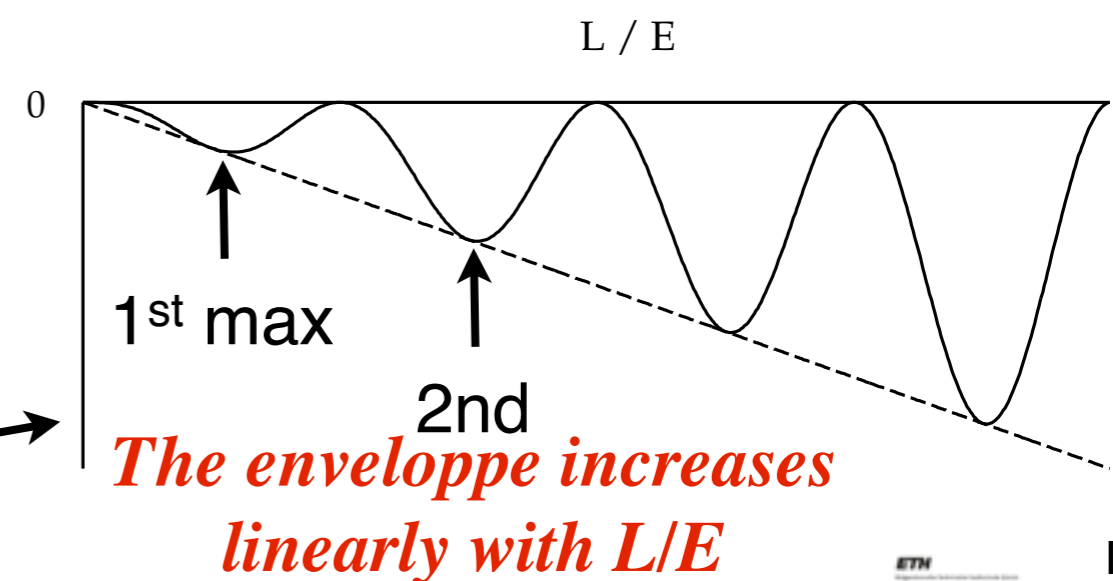
$$a \equiv 2\sqrt{2}G_F n_e E = 7.56 \times 10^{-5} \text{eV}^2 \frac{\rho}{\text{g cm}^{-3}} \frac{E}{\text{GeV}}$$

- Difference between neutrinos and antineutrinos:

$$\begin{aligned}
 \mathcal{A} \equiv P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = & \\
 & 16 \frac{a}{\delta m_{31}^2} \sin^2 \frac{\delta m_{31}^2 L}{4E} c_{13}^2 s_{13}^2 s_{23}^2 (1 - 2s_{13}^2) \\
 & - 4 \frac{aL}{2E} \sin \frac{\delta m_{31}^2 L}{2E} c_{13}^2 s_{13}^2 s_{23}^2 (1 - 2s_{13}^2) \\
 & - 8 \frac{\delta m_{21}^2 L}{2E} \sin^2 \frac{\delta m_{31}^2 L}{4E} s_\delta c_{13}^2 s_{13} c_{23} s_{23} c_{12} s_{12}
 \end{aligned}$$

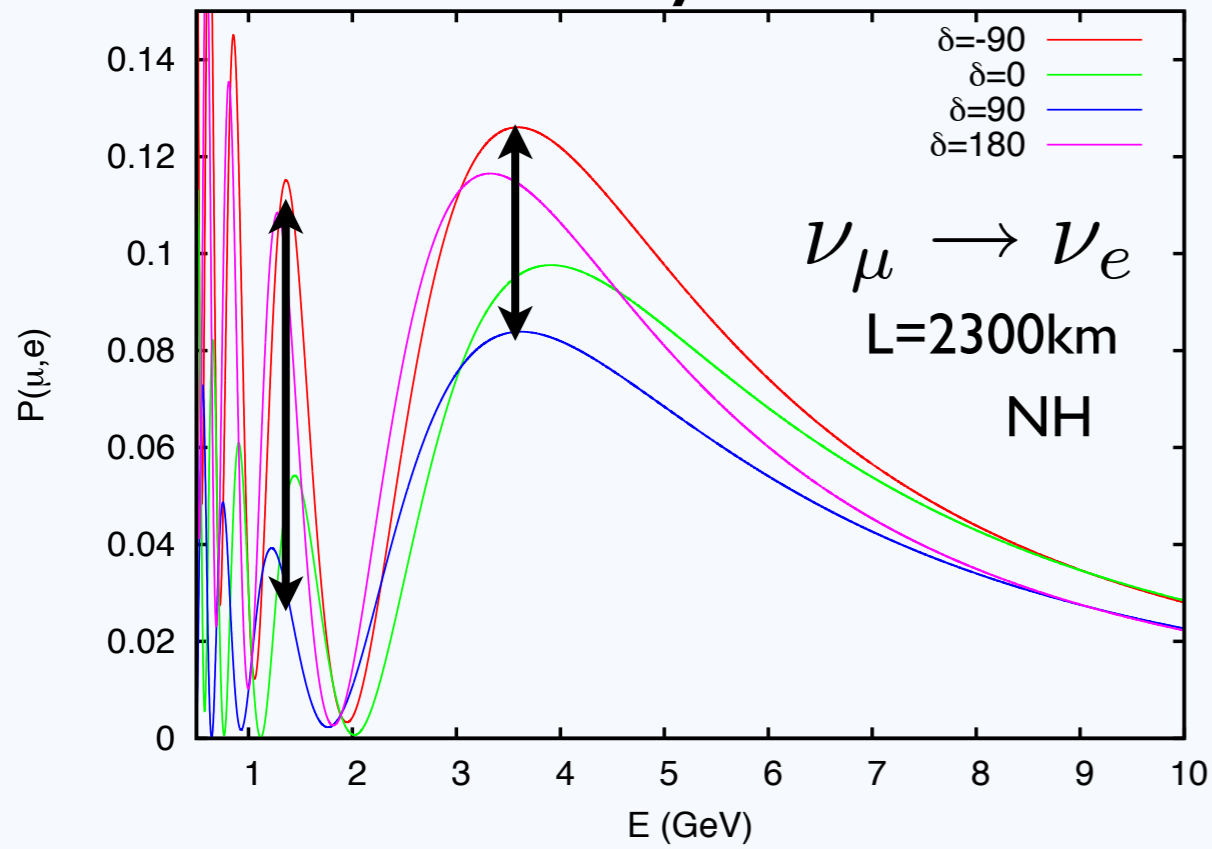
**Matter terms**  
**Pure CP-term**

$$\left. \frac{P(\nu) - P(\bar{\nu})}{P(\nu) + P(\bar{\nu})} \right|_{a=0} \approx -\frac{2s_\delta c_{12} s_{12}}{s_{13}} \cot \theta_{23} \frac{\delta m_{21}^2 L}{2E}$$

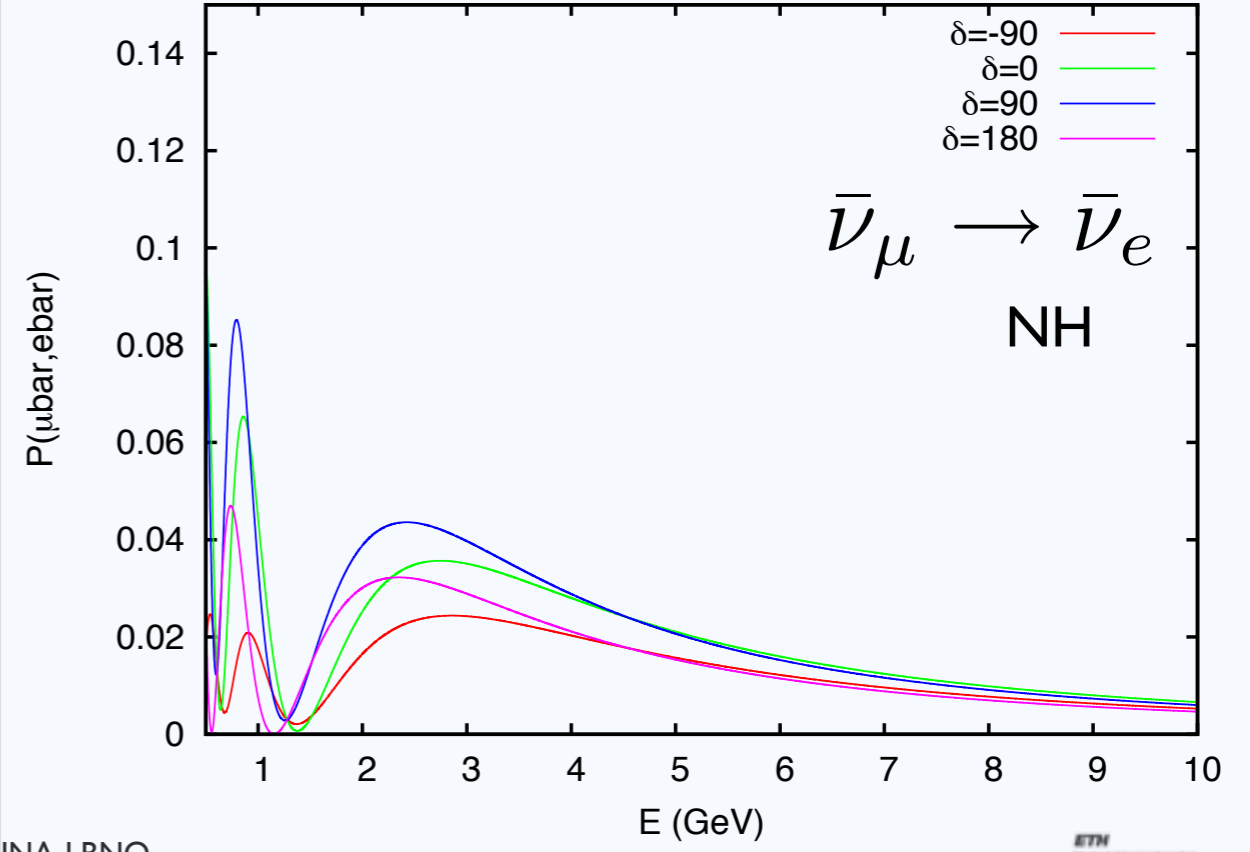
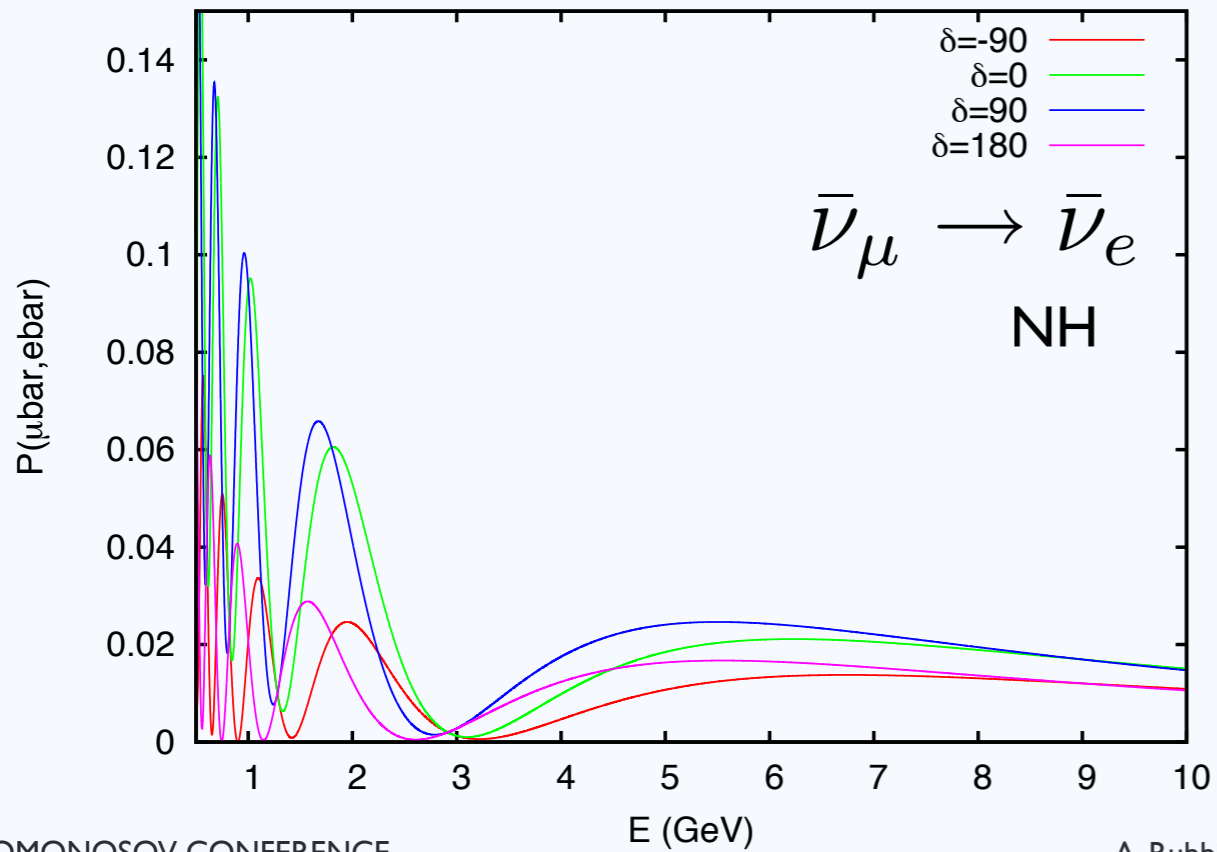
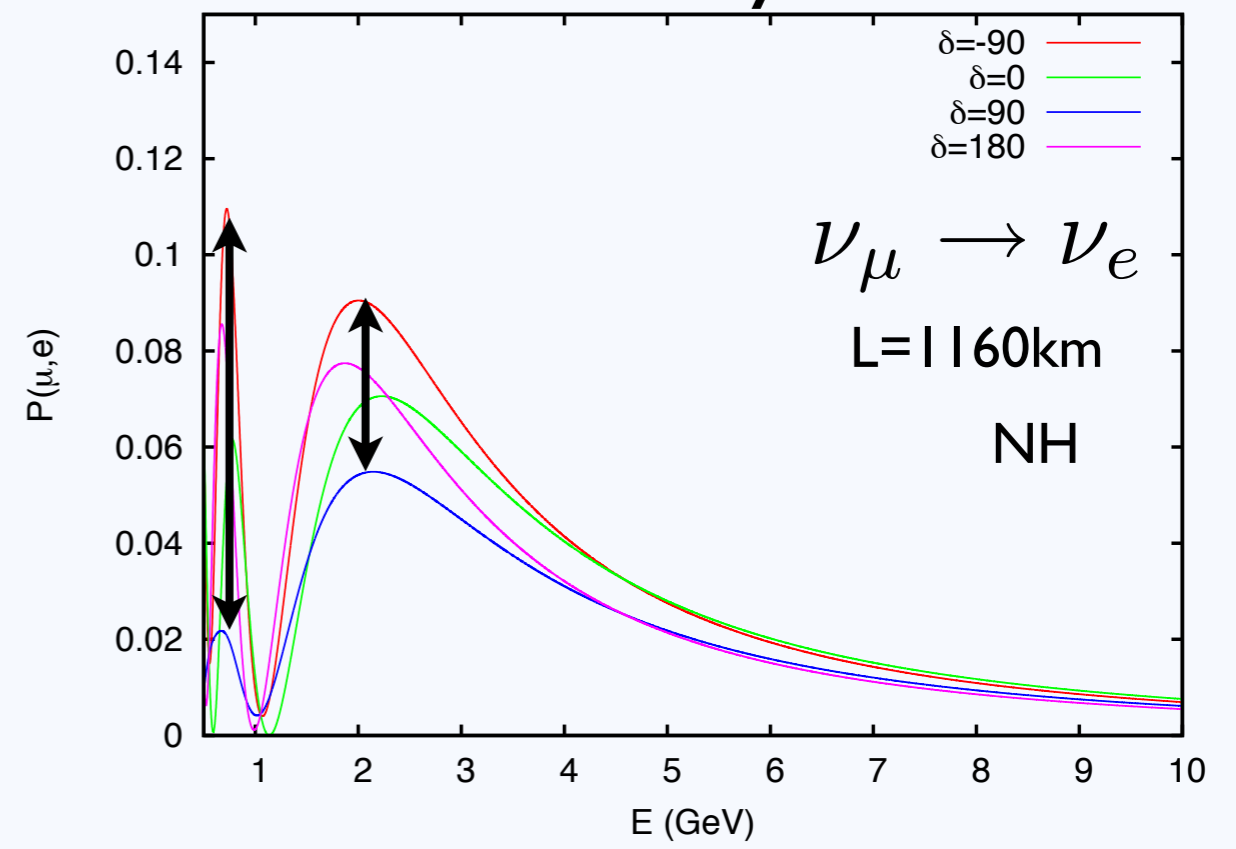


# Expected oscillation probability

## CERN-Pyhäsalmi



## Protvino-Pyhäsalmi



# Event rates/year for 20 kton

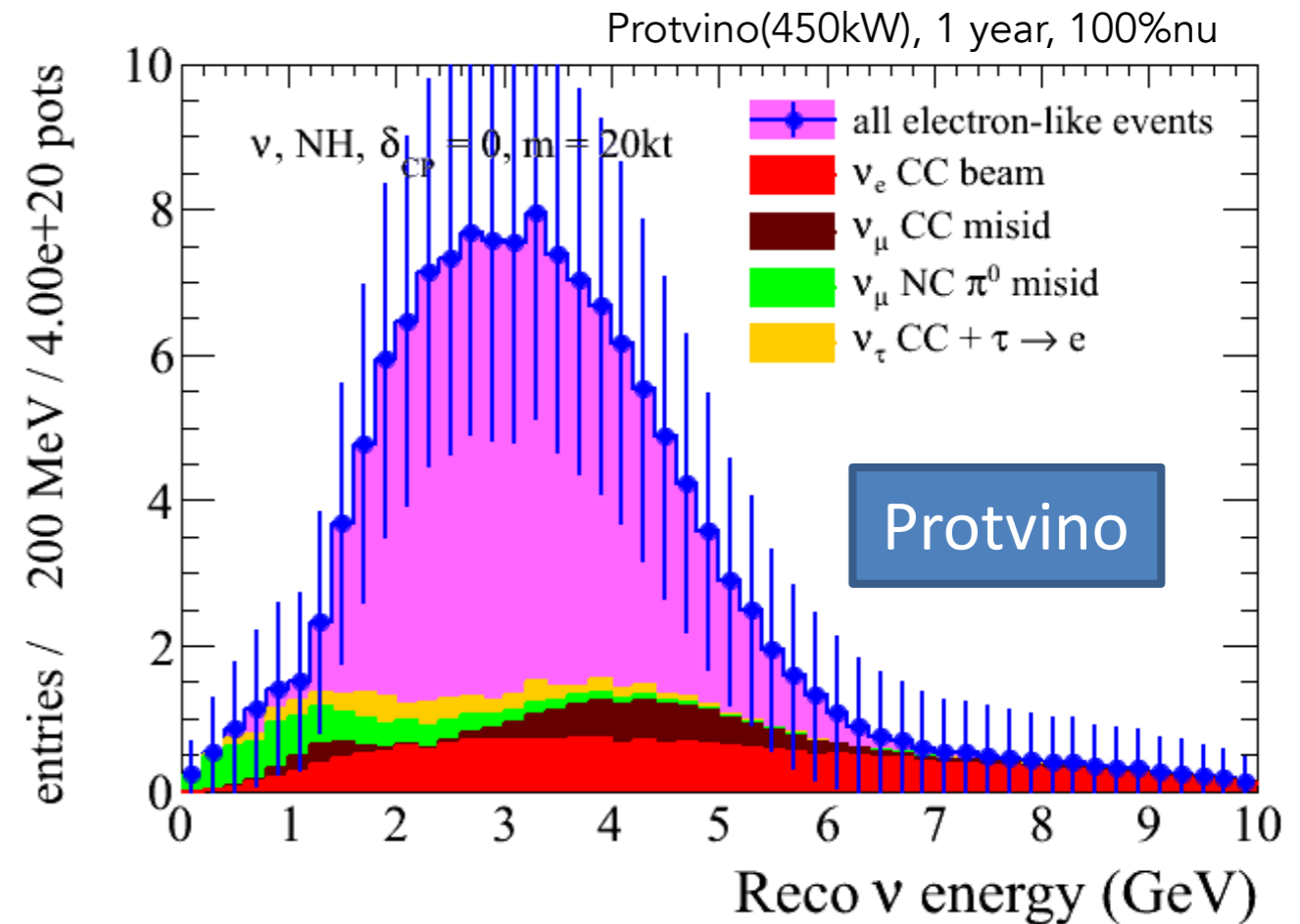
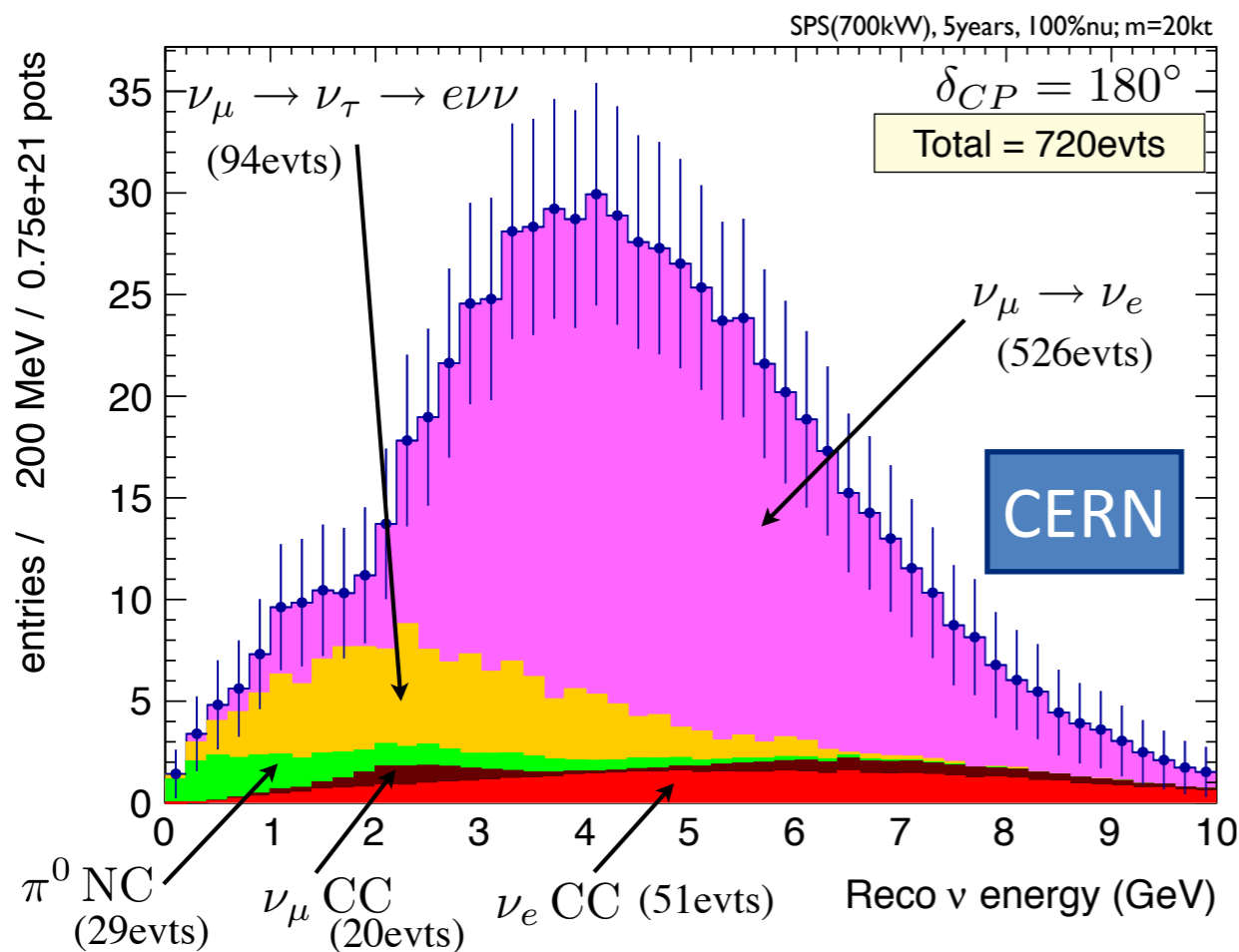


## POT normalisation

Protvino:  $4e20$  pot @ 70 GeV

(corresponds events/1 year): CERN: SPS  $1.5e20$  pot @ 400GeV and HP-PS  $3.5e21$  pot @ 50 GeV

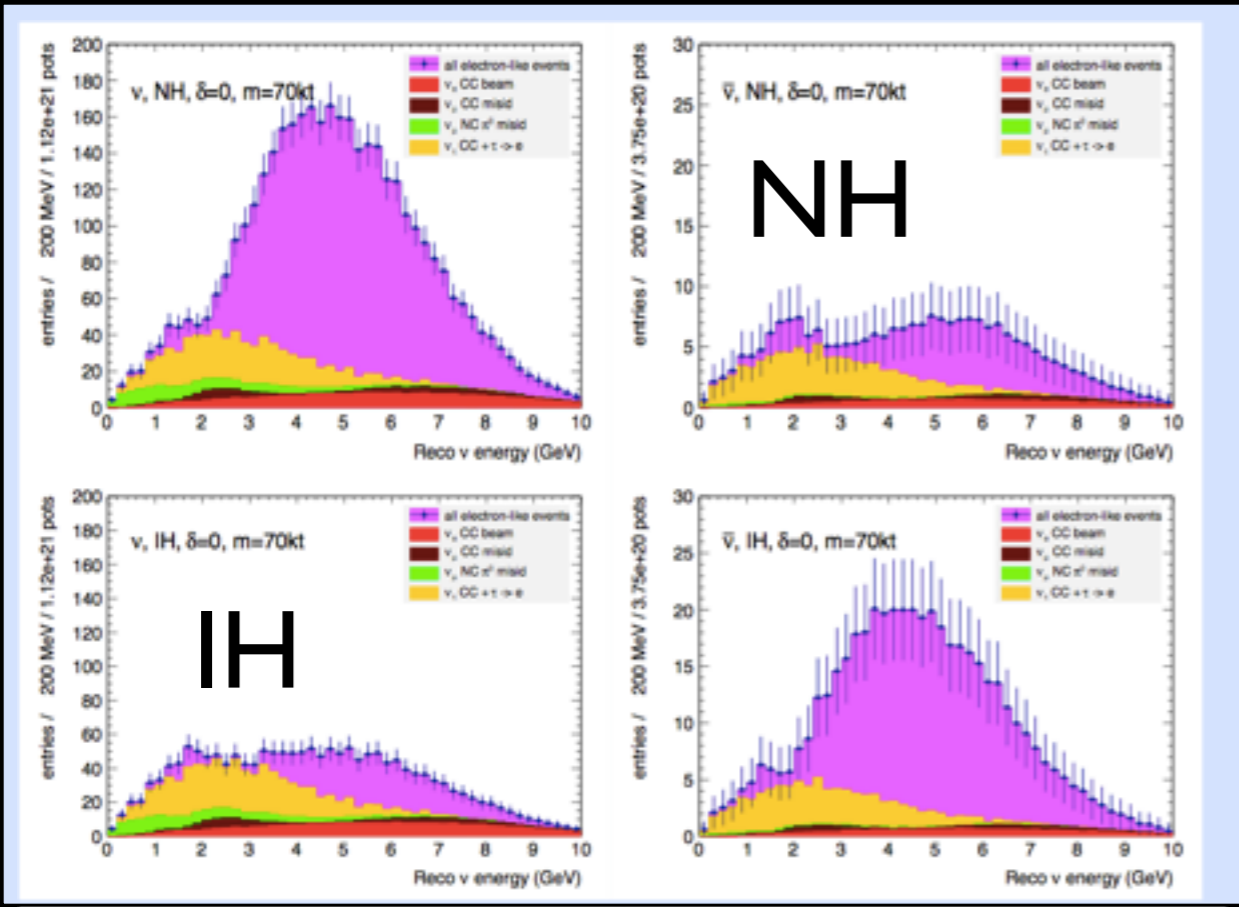
Nu beam	CERN SPS 700kW		CERN HP-PS 2MW		Protvino 450kW	
	$\nu_\mu$ CC	$\nu_e + \bar{\nu}_e$ CC	$\nu_\mu$ CC	$\nu_e + \bar{\nu}_e$ CC	$\nu_\mu$ CC	$\nu_e + \bar{\nu}_e$ CC
NEUT					2056	21
GENIE	<b>1428</b>	10	<b>4007</b>	26	<b>1805</b>	18
GLOBES	1426	10	3975	26	1756	18



# $\delta_{CP}$ & MH dependence

SPS(700kW), 10y, 75%nu-25%antinu; m=70kt

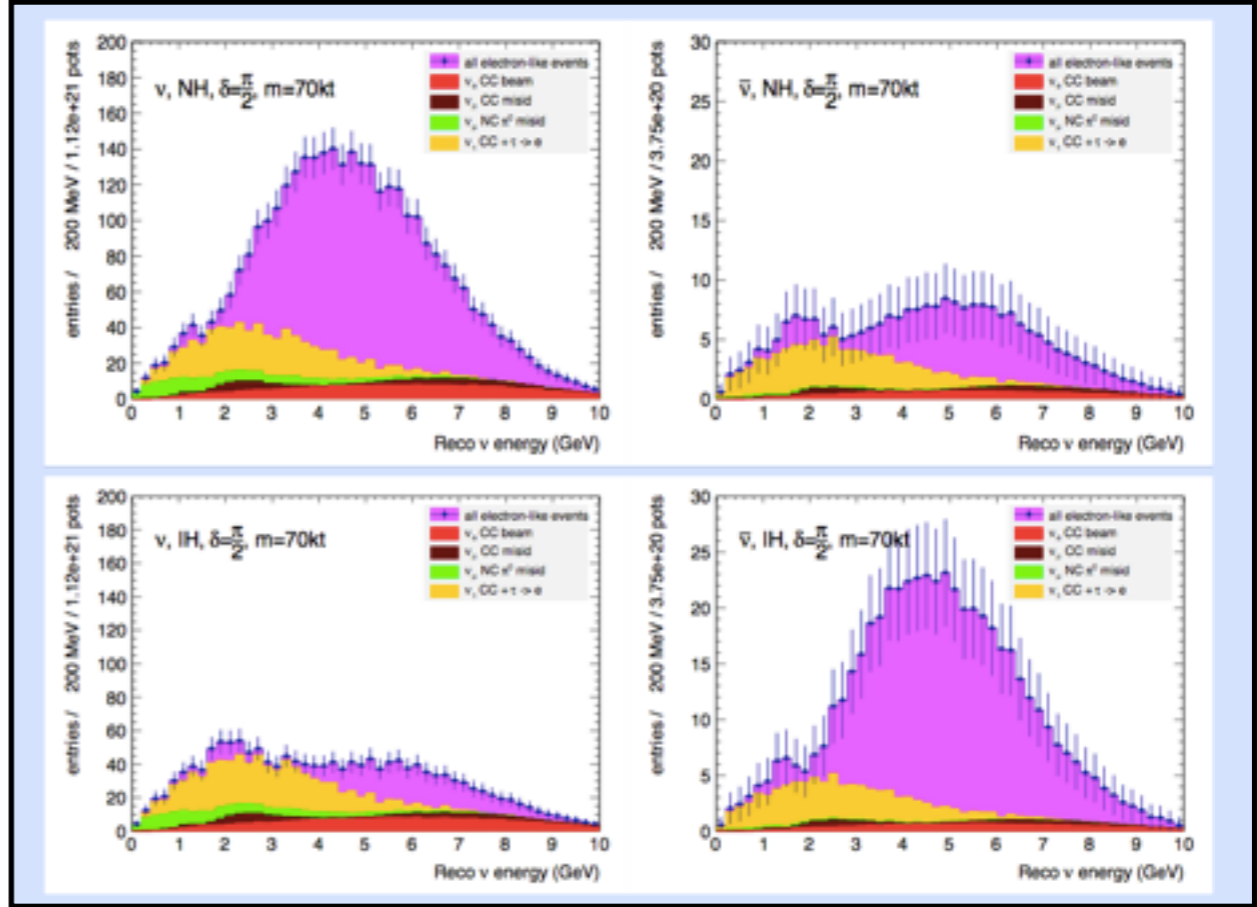
$\delta = 0$



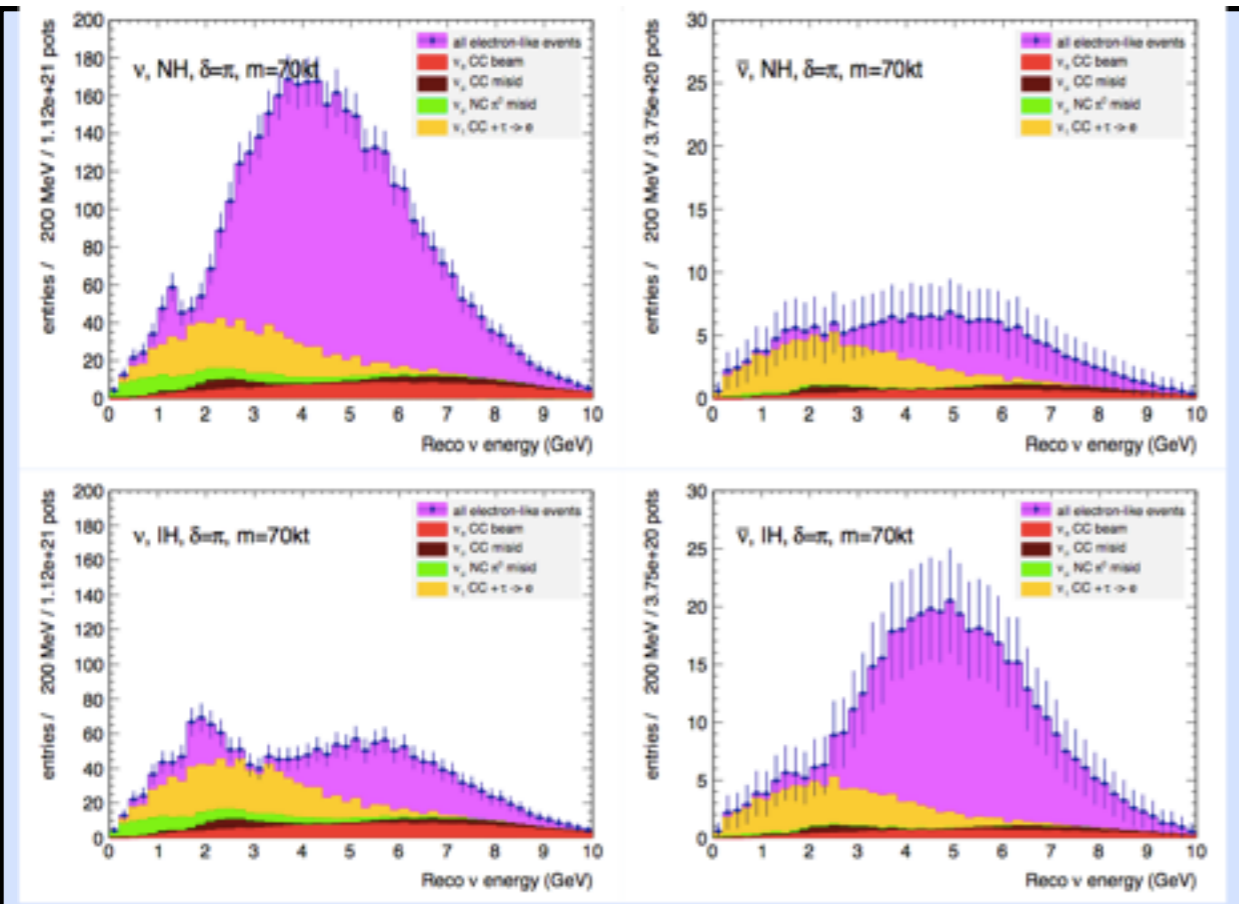
NH

IH

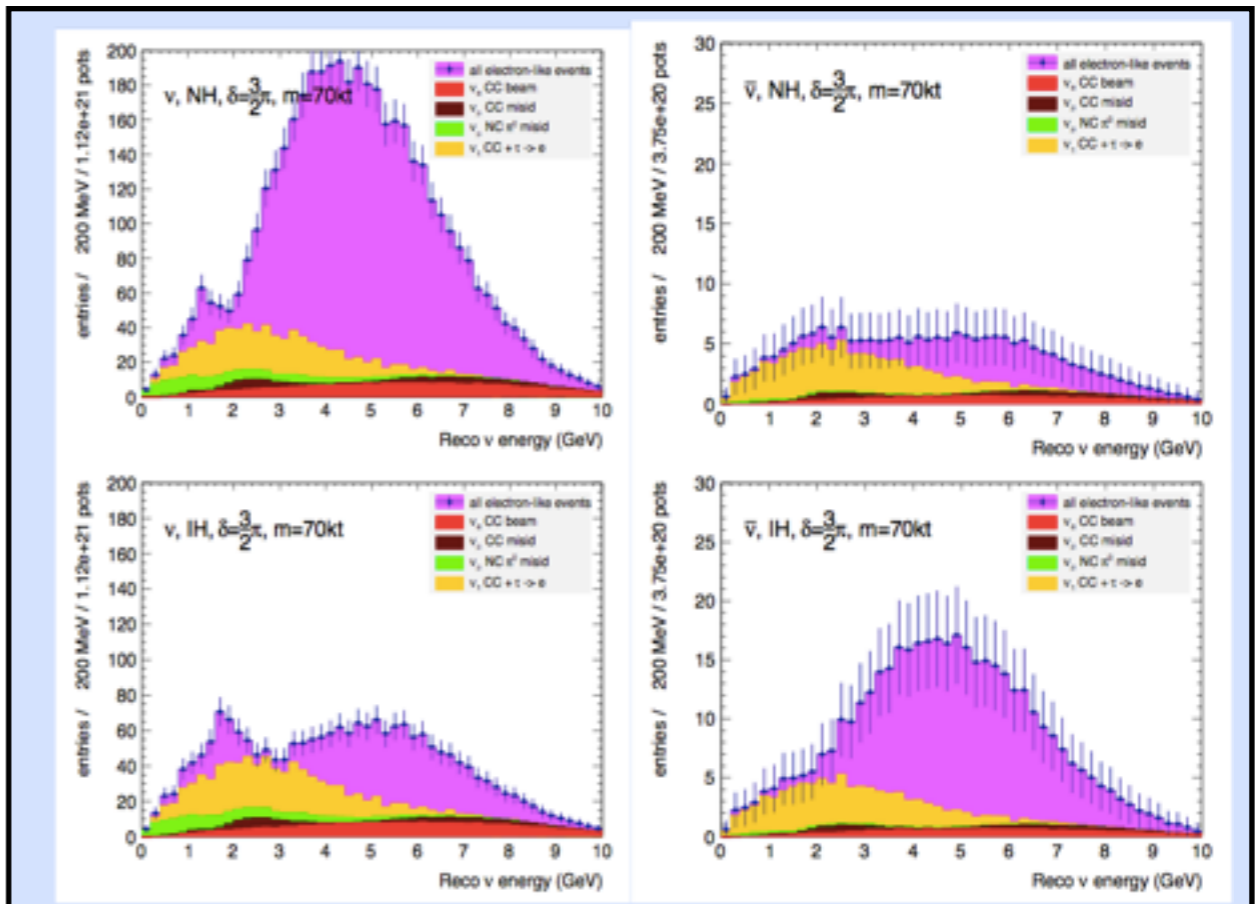
$\delta = \pi/2$



$\delta = \pi$

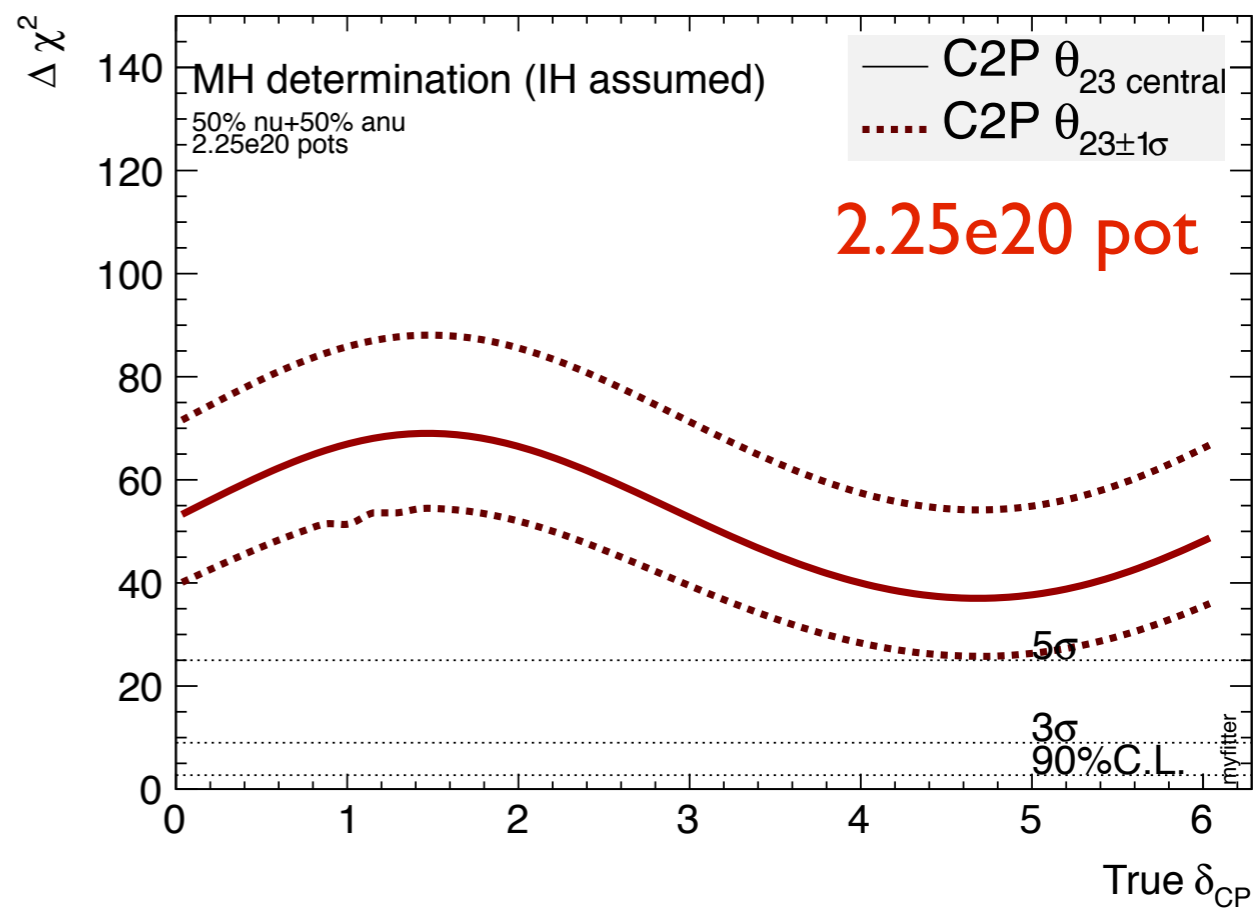
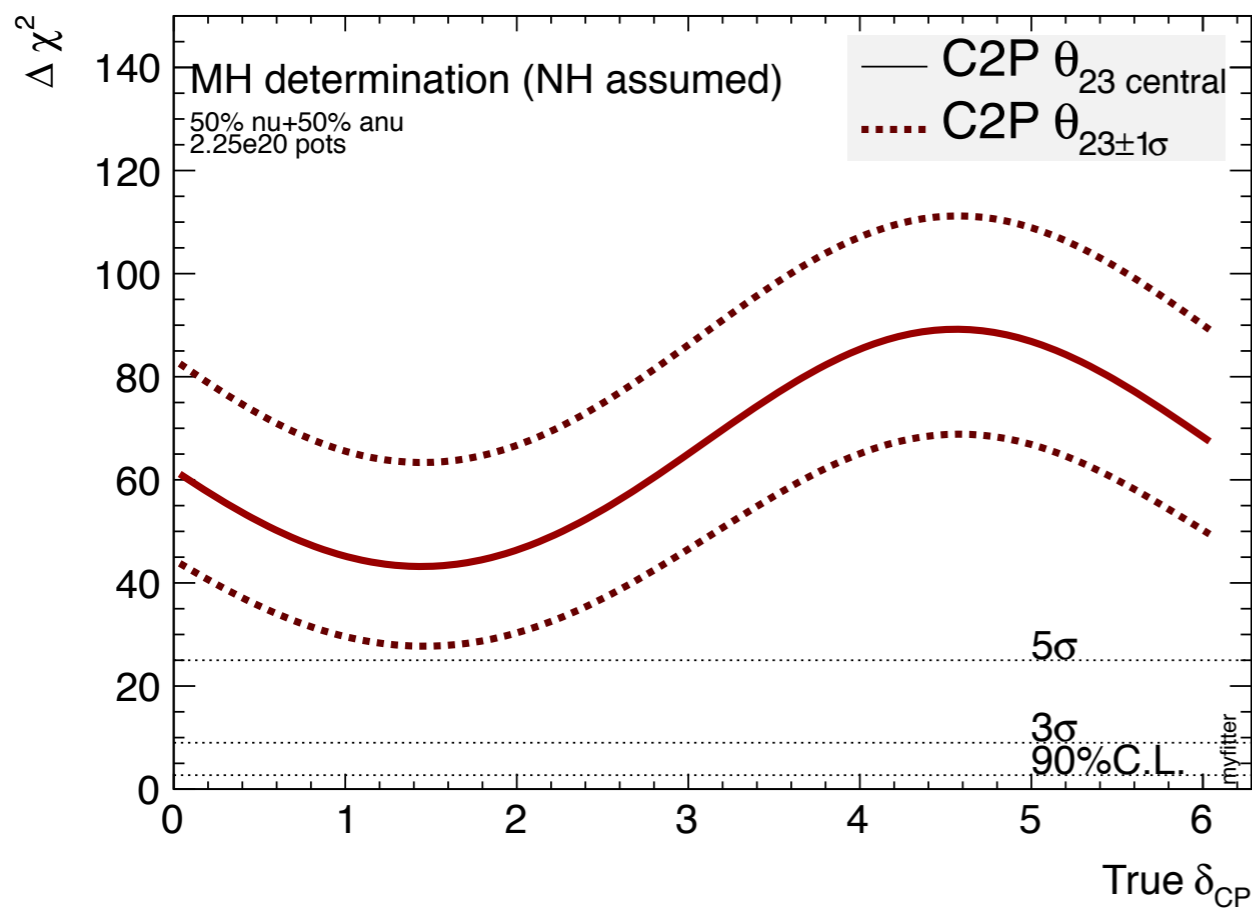


$\delta = 3\pi/2$





**Extracting MH from global fits can not replace a direct  $5\sigma$  measurement from a direct measurement !**



**Provide a  $>5\sigma$  direct determination of MH independent of the values of  $\theta_{23}$  &  $\delta_{CP}$  in  $\approx 2$  years of running**

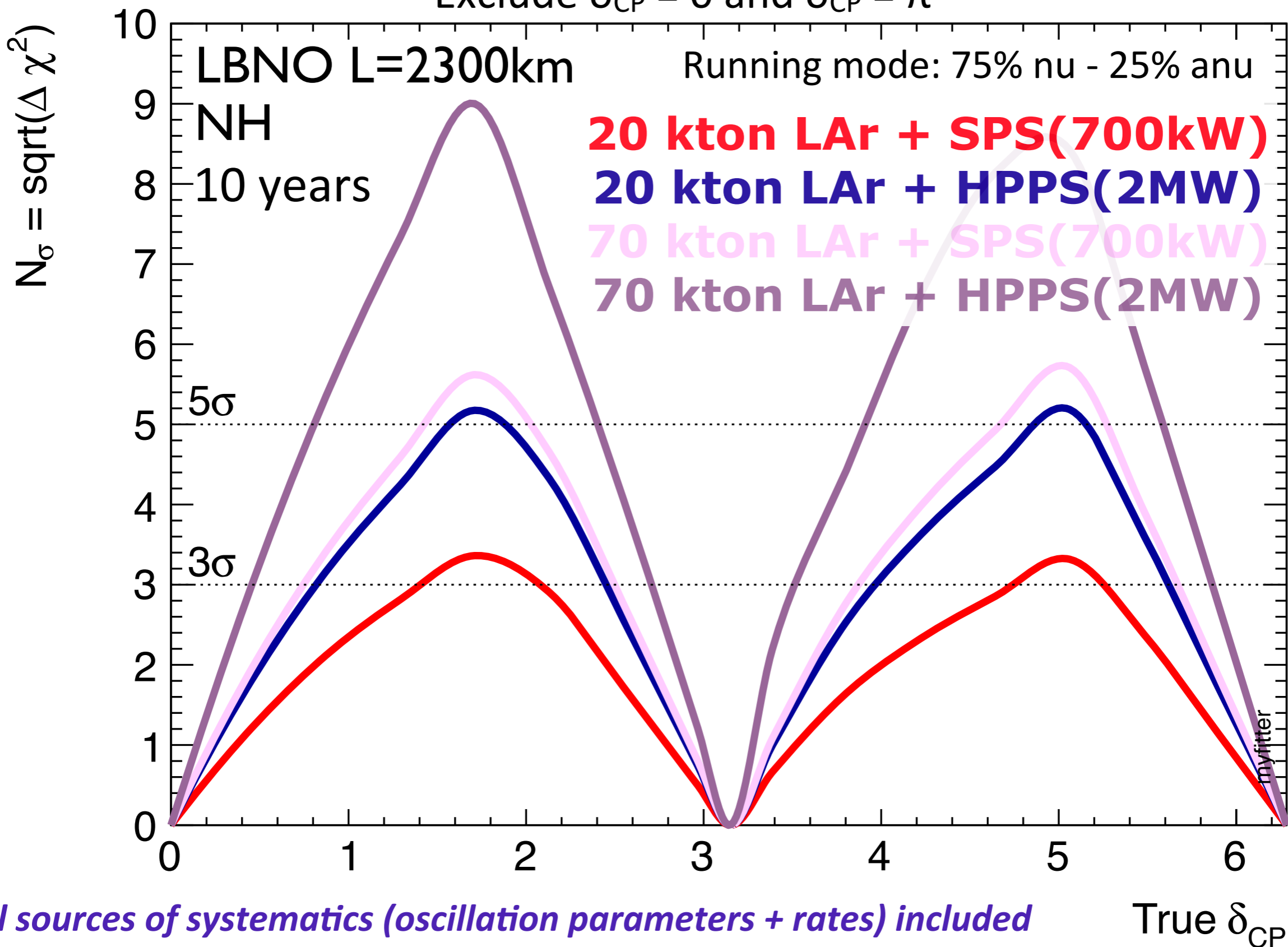
Other methods proposed (atmospheric neutrinos, reactors) do not provide such a level of sensitivity and could be prone to irreducible systematic errors



# Sensitivity to CP violation



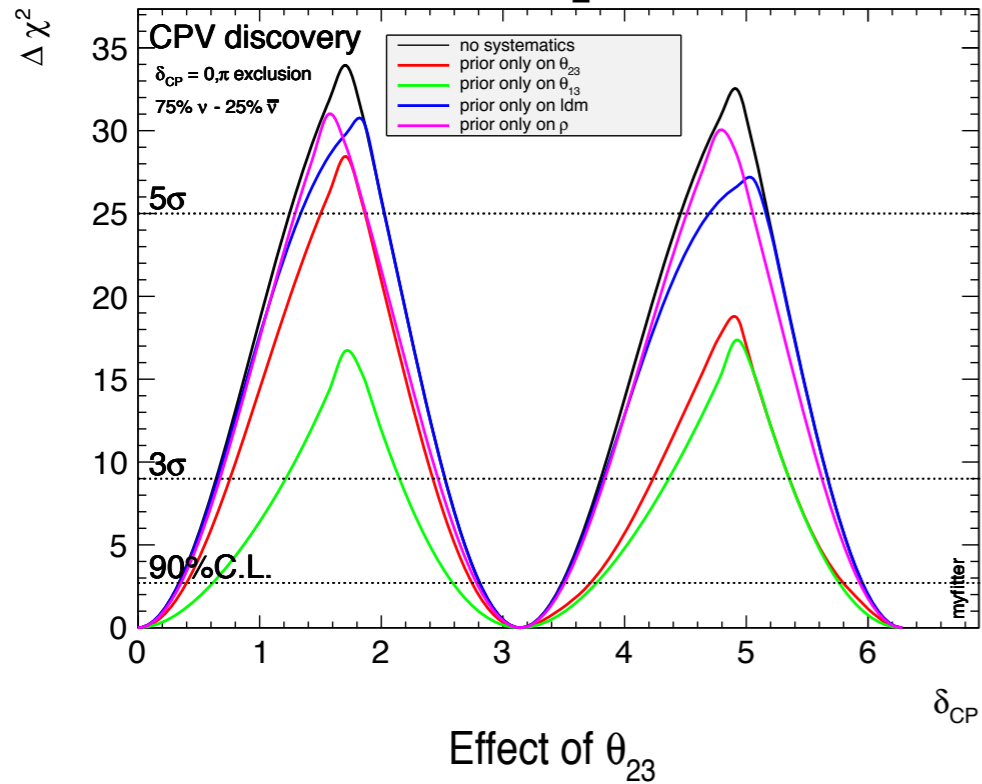
Exclude  $\delta_{CP} = 0$  and  $\delta_{CP} = \pi$



# Effect of systematic errors

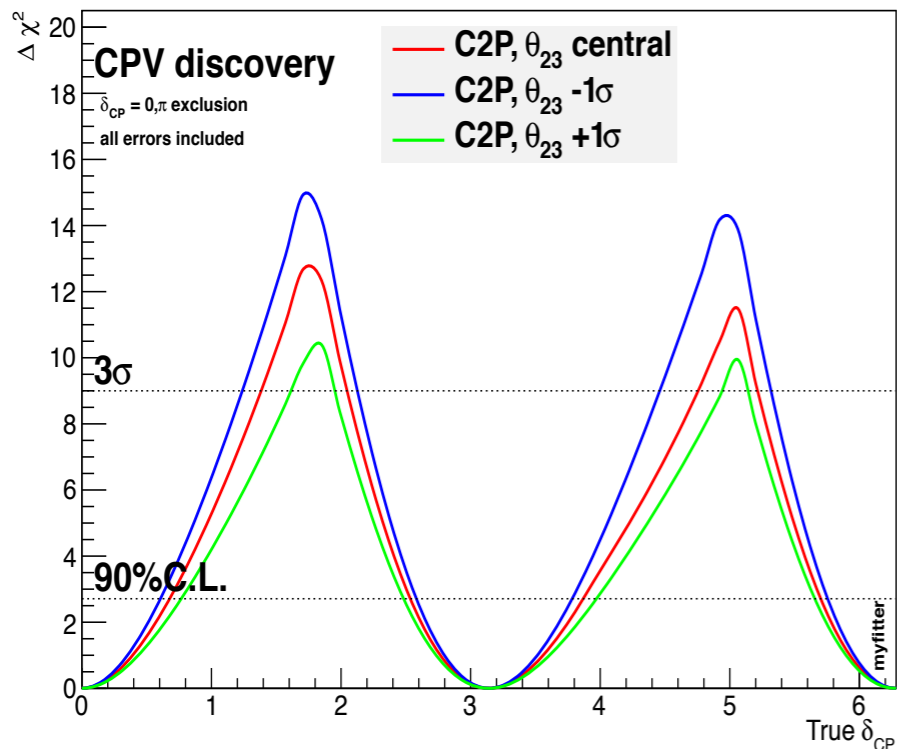
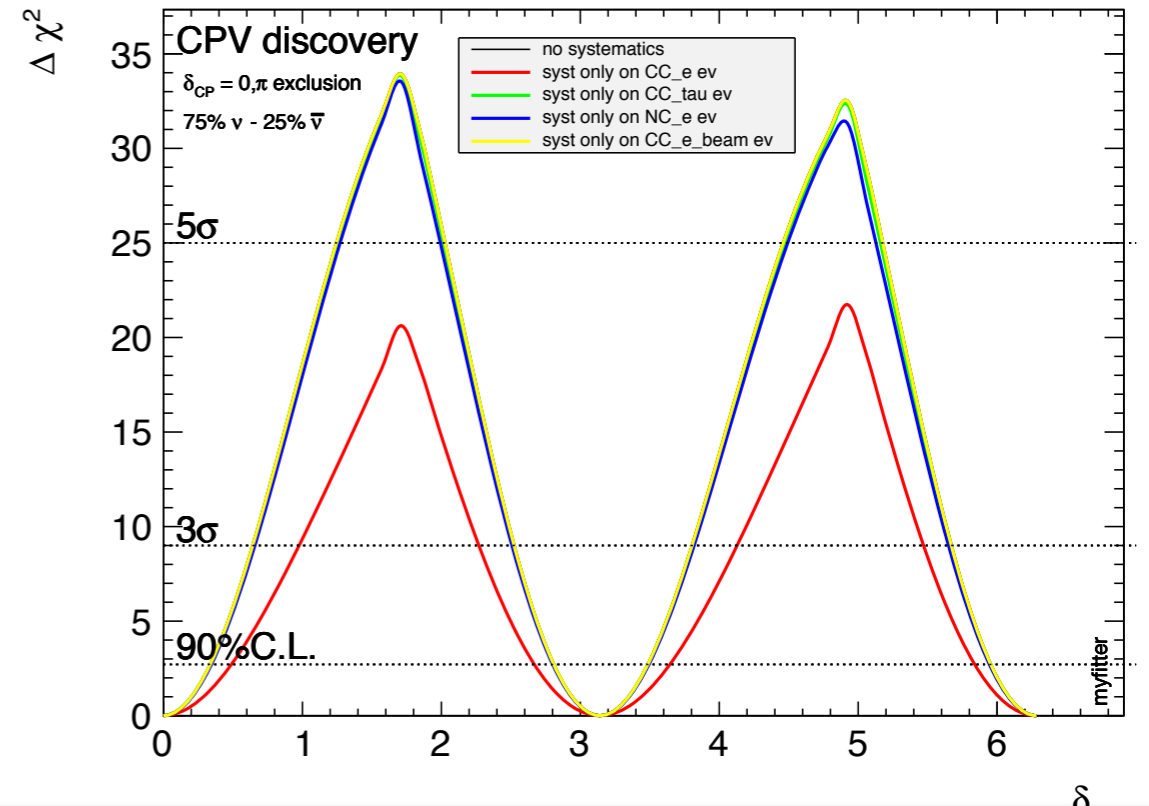
LBNO L=2300km, 20 kton, 10 years

## Oscillation parameters



Effect of  $\theta_{23}$

## Detector related

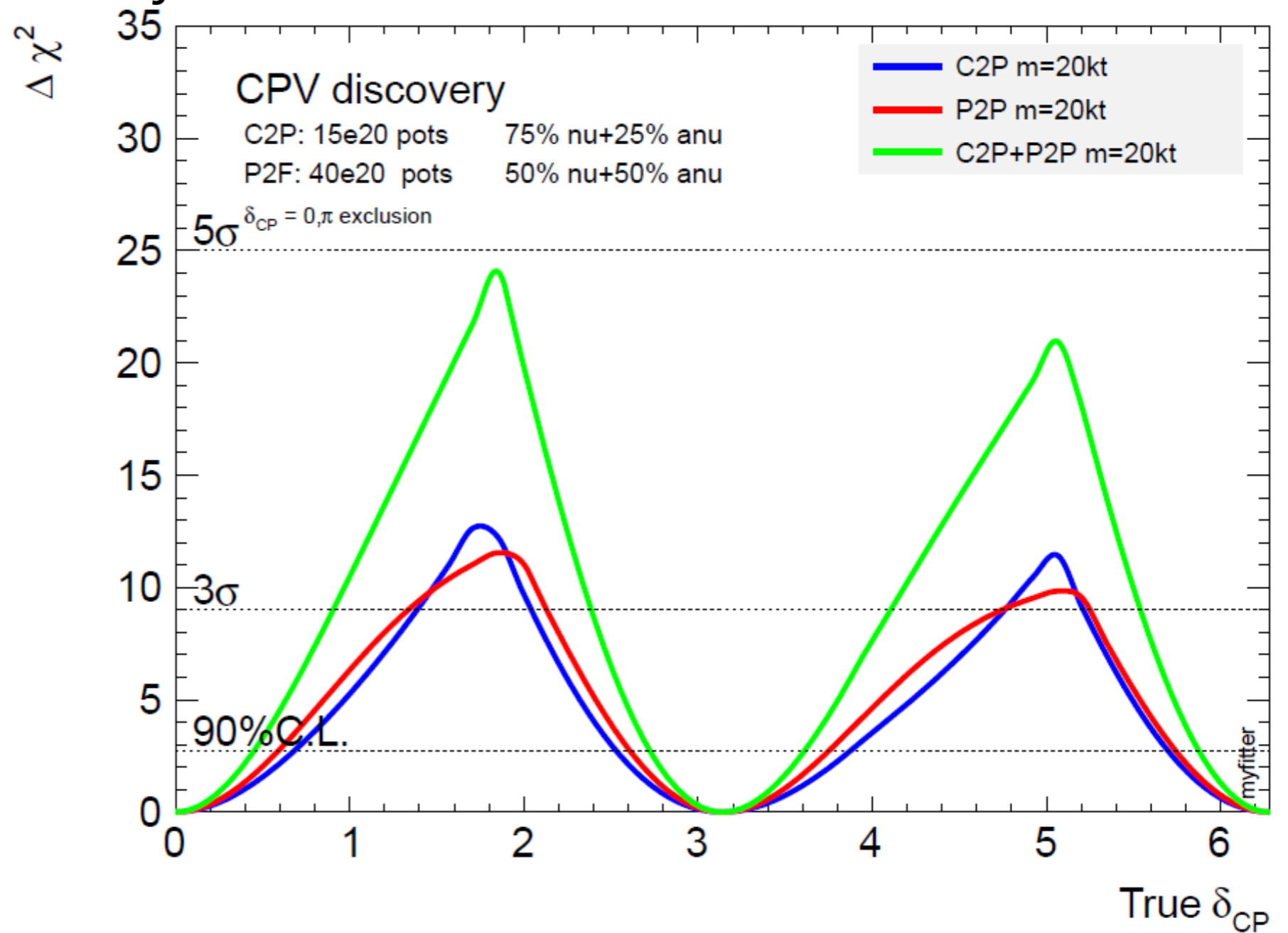


**Without systematic errors, the 20 kton can reach 5sigma CPV in 10 years !**  
**The most important oscillation parameters are  $\theta_{23}$  and  $\theta_{13}$  and the most important systematics is the knowledge of the absolute rate of  $\nu_e$  CC events.**

# LBNO++: two beam experiment!

CERN-Protvino-Pyhäsalmi – 1160&2300 km baselines

study in progress

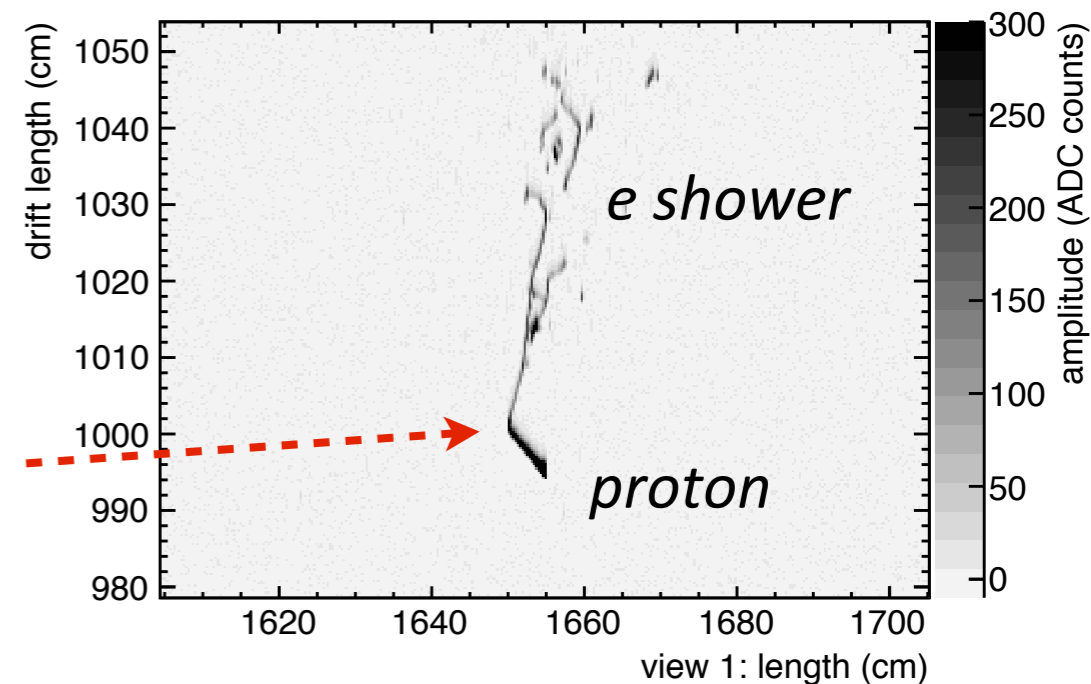


**Sensitivity can be significantly enhanced and systematic errors cross-checked with CERN+Protvino beams with different baselines in the single far LBNO experiment → Unique setup!**

# Atmospheric neutrinos

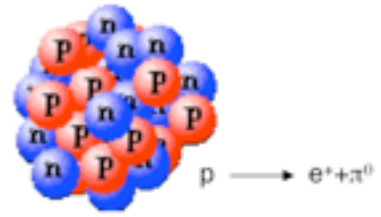
<u>Mode</u>	<u>Events/20kt/yr</u>
$\nu_e$ CC	1440
$\bar{\nu}_e$ CC	310
$\nu_\mu$ CC	2440(w/o osc)
$\bar{\nu}_\mu$ CC	680(w/o osc)
$\nu$ NC	640

MC:  $\nu_e$  CC



- **Neutrino oscillation physics complementary to long baseline beam**
- Clean  $\nu_e$  &  $\nu_\mu$  CC over all range of energies (GeV, MultiGeV)
- Good neutrino energy and angular reconstruction
- Recoil hadronic system on an event-by-event basis
- Statistical separation of  $\nu$  and anti- $\nu$  by exclusive final states
- $\nu_\mu \rightarrow \nu_\tau$  appearance significance  $>3\sigma$  after 3 years exposure ( $\approx 12 \nu_\tau$  CC / year)

# Proton decay sensitivity



For a 20kton exposure of 10 years (200 kton×year)

JHEP 0704 (2007) 041

Mode	Lifetime (90%C.L.)
$p \rightarrow \nu K^+$	$>3 \times 10^{34}$ yrs
$p \rightarrow e^+ \gamma, p \rightarrow \mu^+ \gamma$	$>3 \times 10^{34}$ yrs
$p \rightarrow \mu^- \pi^+ K^+$	$>3 \times 10^{34}$ yrs
$n \rightarrow e^- K^+$	$>3 \times 10^{34}$ yrs
$p \rightarrow \mu^+ K^0, p \rightarrow e^+ K^0$	$>1 \times 10^{34}$ yrs
$p \rightarrow e^+ \pi^0$	$>1 \times 10^{34}$ yrs
$p \rightarrow \mu^+ \pi^0$	$>0.8 \times 10^{34}$ yrs
$n \rightarrow e^+ \pi^-$	$>0.8 \times 10^{34}$ yrs

Expect  $\approx$ linear sensitivity improvement with exposure until 1000 kton×year

# Supernova detection channels



JCAP 0310 (2003) 009

JCAP 0408 (2004) 001

For 20 kton and a SN explosion  
at the distance of 5 kpc:

$$\langle E_{\nu_e} \rangle = 11\text{MeV}, \langle E_{\bar{\nu}_e} \rangle = 16\text{MeV}, \langle E_{\nu_x} \rangle = \langle E_{\bar{\nu}_x} \rangle = 25\text{MeV}$$

**Events:**

$$\nu_e \text{ } ^{40}\text{Ar} \rightarrow e^- \text{ } ^{40}\text{K}^* \quad (E_\nu > 1.5 \text{ MeV}) \quad \approx 23820$$

$$\bar{\nu}_e \text{ } ^{40}\text{Ar} \rightarrow e^+ \text{ } ^{40}\text{Cl}^* \quad (E_\nu > 7.48 \text{ MeV}) \quad \approx 2420$$

$$\nu_x \text{ } ^{40}\text{Ar} \rightarrow \nu_x + \text{ } ^{40}\text{Ar}^* \quad \approx 30440$$

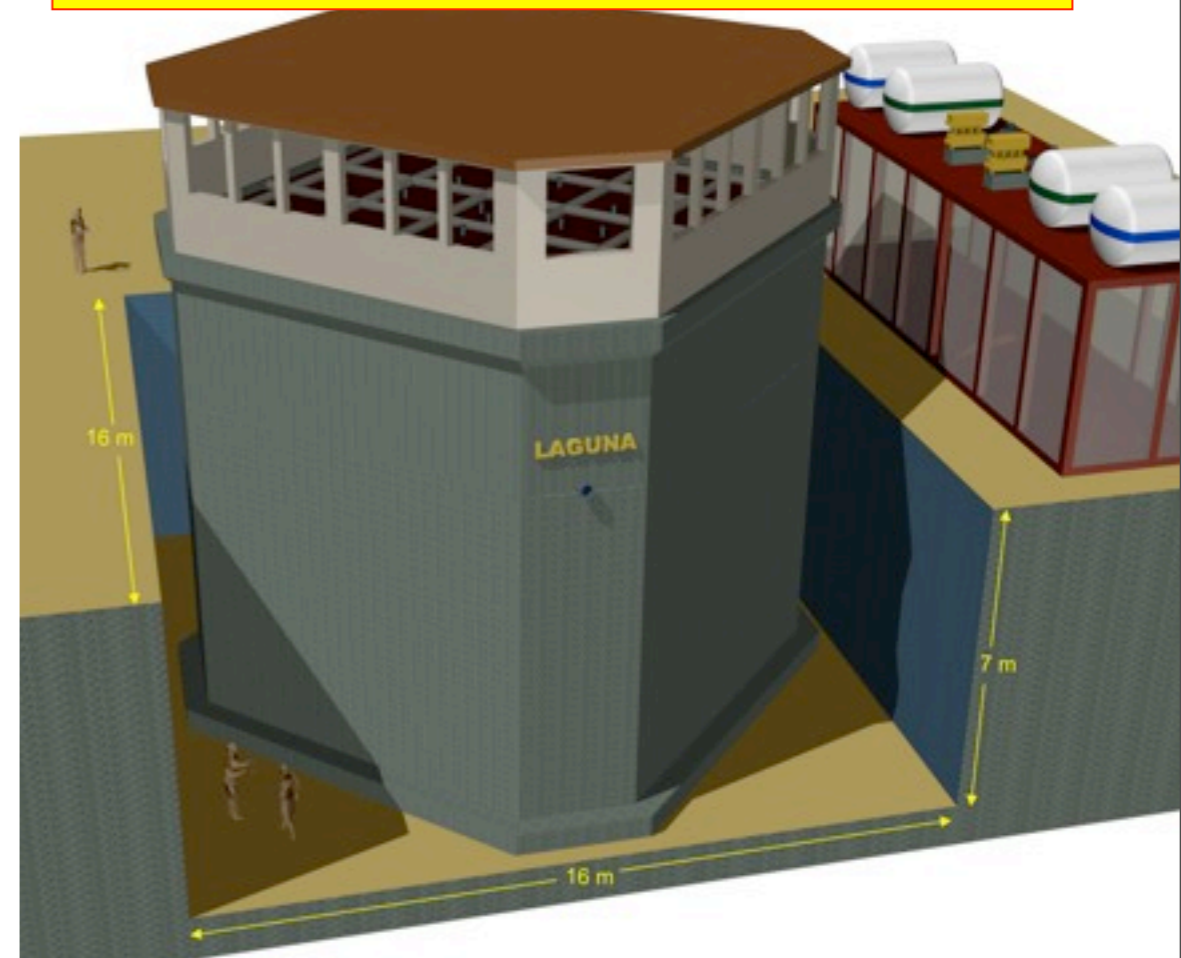
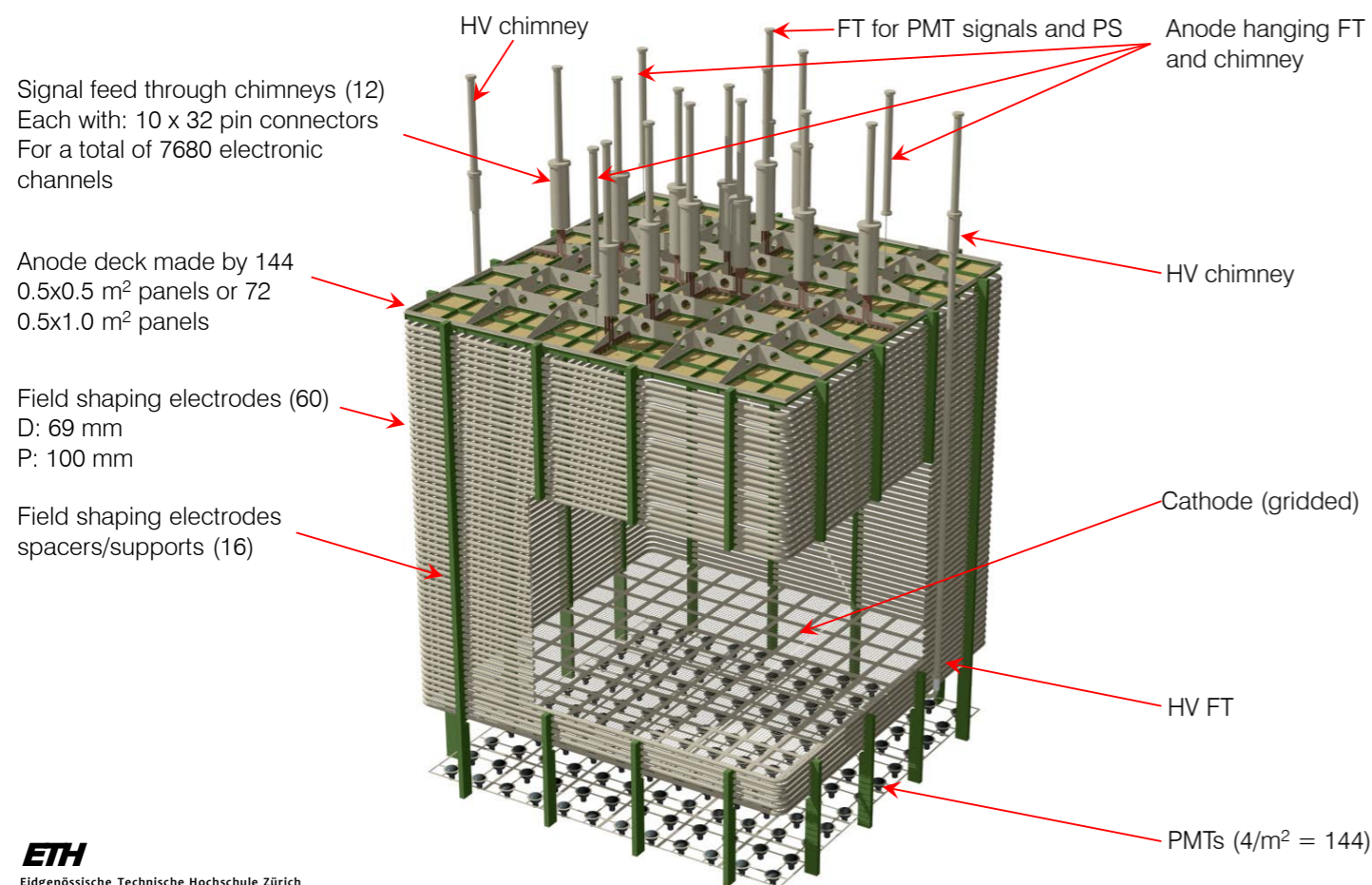
$$\nu_x e^- \rightarrow \nu_x e^- \quad \approx 1330$$

- Unique sensitivity to electron neutrino flavour (most other SN-detectors detect inverse beta decays)
- Combined analysis of all reaction modes
- Neutrino mass via TOF

# Double phase LBNO LAr demonstrator

- We are proposing a **6x6x6 = 216 m<sup>3</sup> active volume double phase LAr detector** to be constructed and operated at CERN
- Charged test beams to collect the large controlled data set allowing **electromagnetic and hadronic calorimetry** and general **detector performance** (PID, ...) to be measured, **simulation and reconstruction** to be improved and validated.
- Proposal to SPSC ready for submission

CERN committed to provide support for these Important developments. FNAL interest.





# Conclusion

- EU FP7 LAGUNA/LBNO design study has made significant progress at designing and optimising a next generation deep underground neutrino observatory.
- LBNO (SPSC-EOI-007) has been put forward to CERN with unique physics potentials, including astro-particle physics and proton decay search.
- It is conceived as an incremental approach starting with an underground LAr detector, a clear stage 1 physics goal and well-defined upgrade plan, in order to reach a CPV discovery.
- Protvino could complement LBNO, greatly enhancing its physics potential. A Protvino near-detector complex can provide important neutrino physics.
- We are now proposing to CERN an LBNO demonstrator for the double phase LAr technology at a relevant scale (216m<sup>3</sup>).
- The LAGUNA/LBNO consortium plans to finalise its findings and submit its reports by the end of 2014.

# Acknowledgements

- FP7 Research Infrastructure “Design Studies” LAGUNA (Grant Agreement No. 212343 FP7-INFRA-2007-1) and LAGUNA-LBNO (Grant Agreement No. 284518 FP7-INFRA-2011-1)
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