

# ANTARES and NEMO – toward a $\text{km}^3$ Neutrino Telescope in the Mediterranean Sea

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# OUTLINE

## ► Introduction

## ► The ANTARES and NEMO project

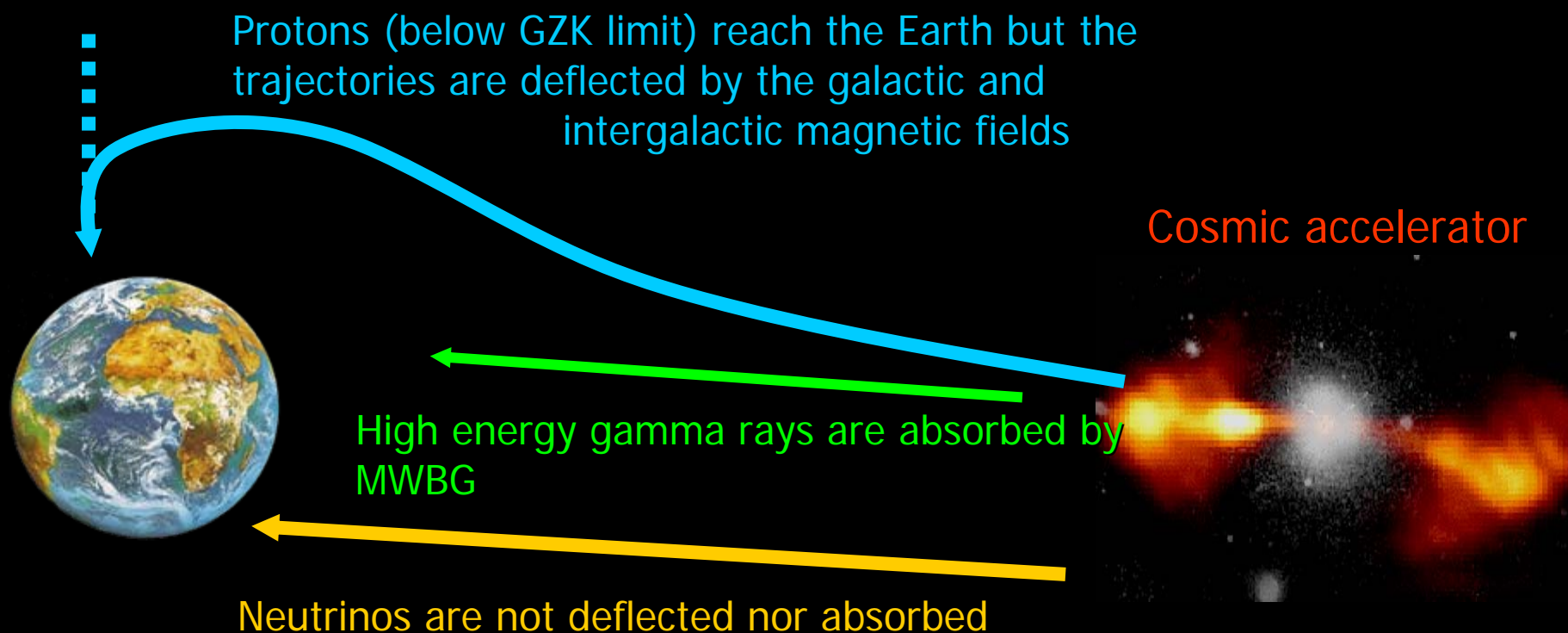
- Description
- Results
- Perspectives

## ► Conclusions

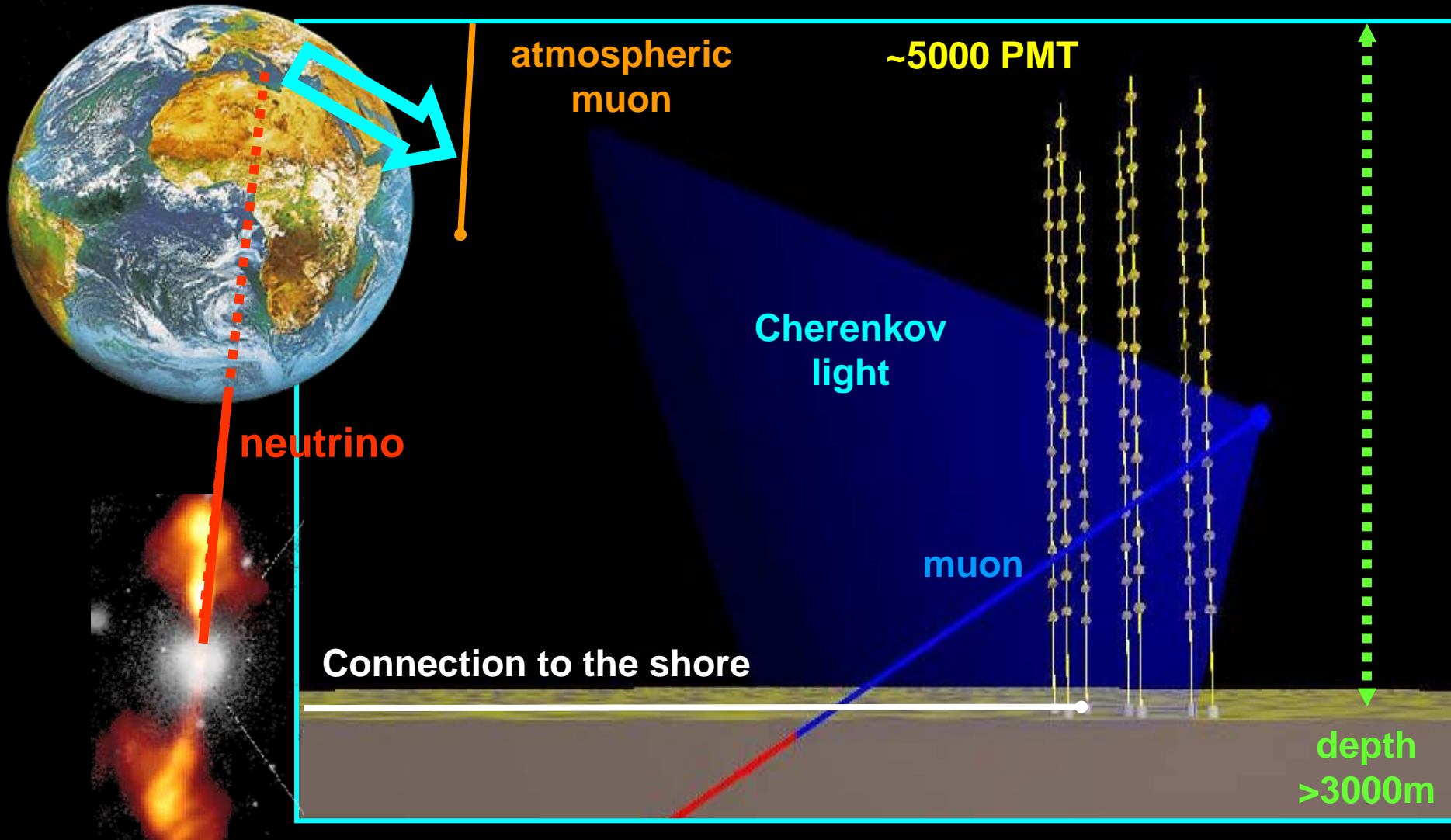
# Physics Motivations

- High Energy Neutrinos are a powerful tool to investigate the hadronic mechanisms of Cosmic Ray production

Proton apparent direction

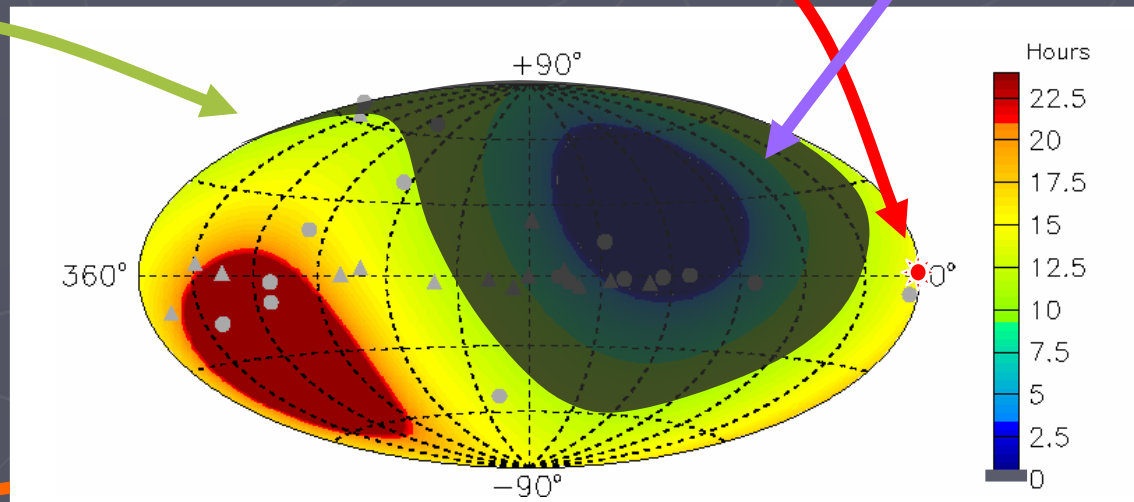


# Physics Motivations



# Physics Motivations

- ▶ A neutrino telescope in the Northern Hemisphere will provide:
  - Complementary sky survey to ICECUBE
  - Overlap observation region with ICECUBE
  - Study of the Galactic Center



# Physics Motivation

- ▶ The small cross section and the expected low neutrino flux require
  - large volume telescope  $\sim 1 \text{ km}^3$
  - long observation time  $\sim \text{yrs}$
- ▶ The atmospheric muon background requires
  - a shielding  $> 2000 \text{ m}$  water equivalent
- ▶ The Mediterranean Sea provides convenient sites

# The Sites

KM3  
Mediterraneo

Pylos



La Seyne



Capo Passero



BAIKAL

DUMAND

• ICECUBE





# The ANTARES Collaboration

## ► Physics Institutes from:

- France
- Germany
- Italy
- NL
- Spain
- Russia
- UK

NEMO

## ► Sea biologists, geophysicists



# The NEMO Collaboration



INFN

- Bari, Bologna, Catania, Genova, LNF, LNS, Napoli, Pisa, Roma

## ► Universities

- Bari, Bologna, Catania, Genova, Napoli, Pisa, Roma "La Sapienza"



CNR

- Istituto di Oceanografia Fisica, La Spezia
- Istituto di Biologia del Mare, Venezia
- Istituto Sperimentale Talassografico, Messina



Istituto Nazionale di Geofisica e Vulcanologia (INGV)



Istituto Nazionale di Oceanografia e Geofisica Sperimentale (OGS)



Istituto Superiore delle Comunicazioni e delle Tecnologie dell'Informazione (ISCTI)

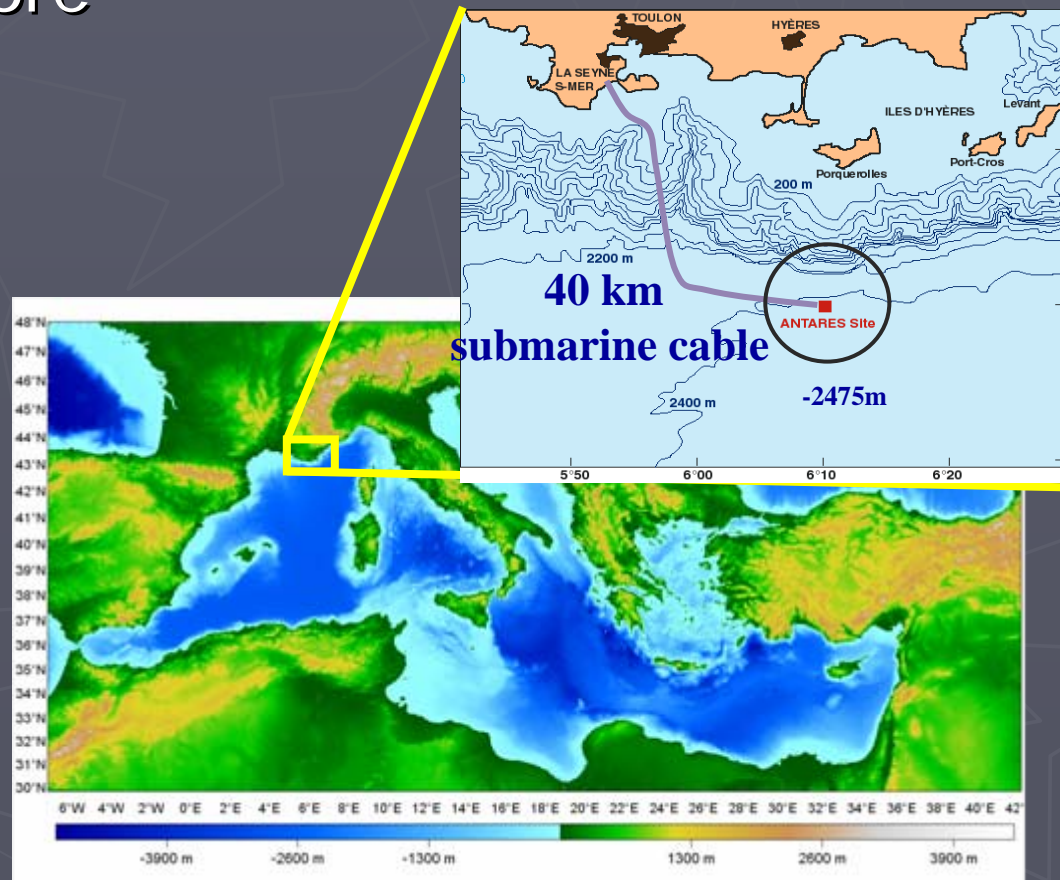
more than 70 researchers involved

# Technical Requirements

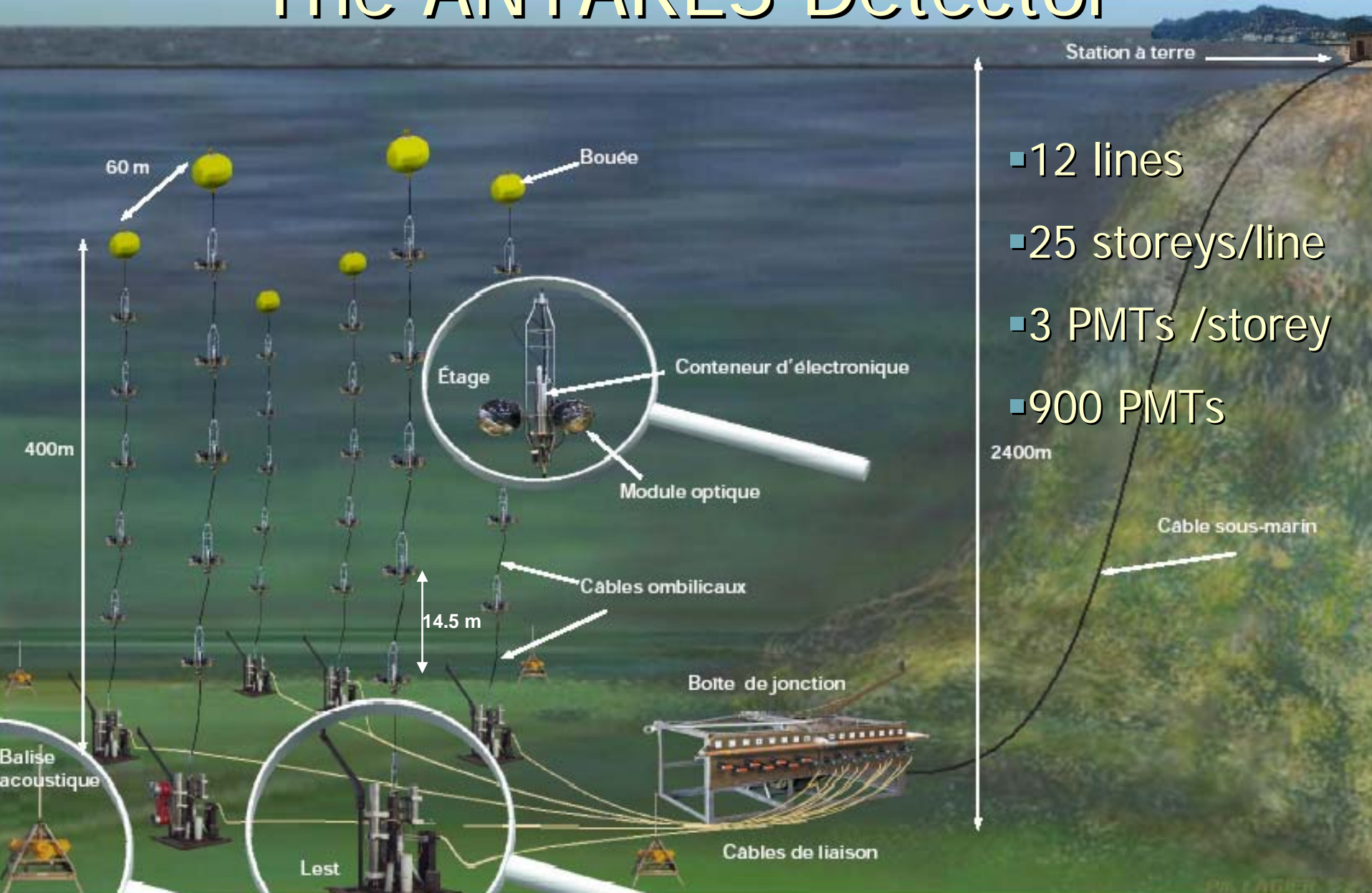
- Issues to be solved to realize a km<sup>3</sup> neutrino telescope in the Mediterranean Sea
  - Best site location (i.e. depth, water quality ...)
  - Optical background from <sup>40</sup>K-decay and bioluminescence
  - Detector deployment
  - Detector rigidity against sea currents
  - Resistance to corrosion from salt water

# The ANTARES Site

- ▶ 2400 m sea depth
- ▶ short path to the shore (40 Km)
- ▶ infrastructures (IFREMER,...)
- ▶ latitude :  $42^{\circ} 50' \text{ N}$



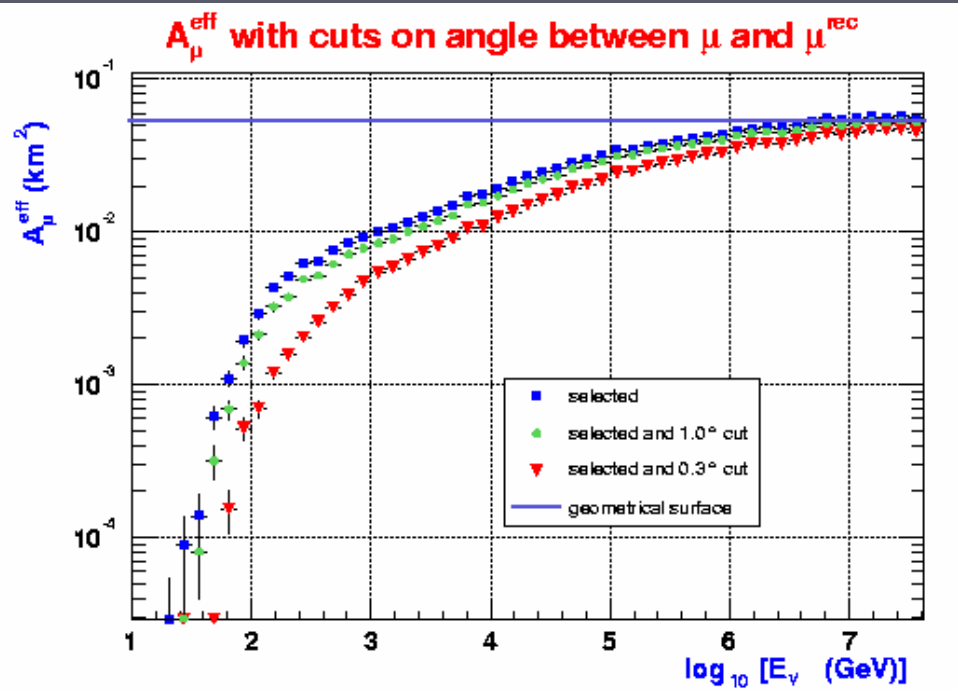
# The ANTARES Detector



- 12 lines
- 25 storeys/line
- 3 PMTs /storey
- 900 PMTs



# ANTARES Expected Performance



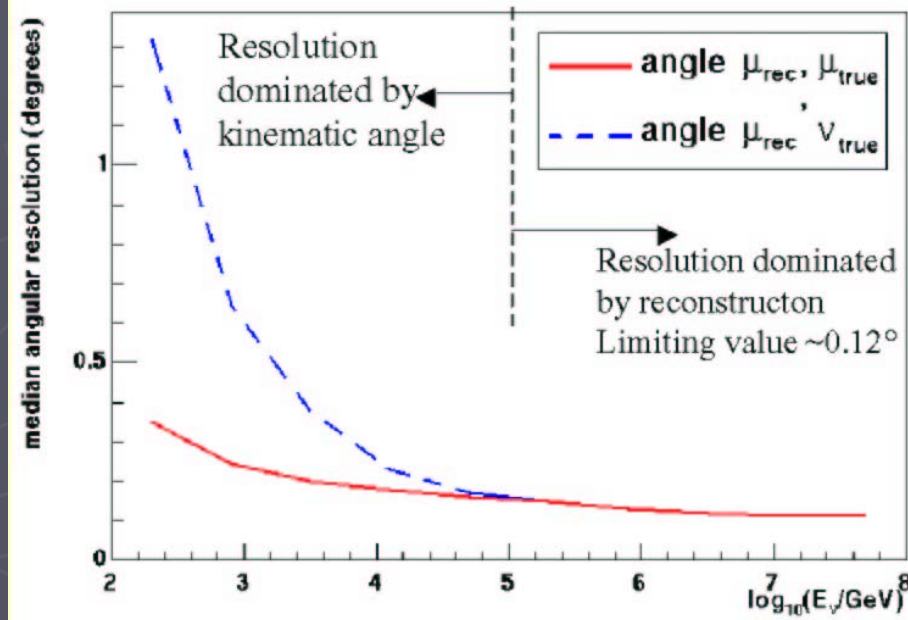
## Angular resolution

- below 10 TeV: dominated by  $\nu_{\mu}$ - $\mu$  angle
- above 10 TeV:  $\leq 0.2^{\circ}$  after reconstruction

## Effective area

depends on :

- reconstruction efficiency
- selection cuts
- absorption length



# ANTARES – Sea Operations

Oct 01: main Electro Optical Cable

Dec 02: Junction Box

Dec 02: Pre-Deployment (PSL)

Feb 03: Mini Instrumentation Line (MIL)



# The ANTARES Project

- ▶ Mar 01: Sea bed survey
- ▶ Oct 01: Electro Optical cable deployment
- ▶ Dec 02: Junction Box (JB) deployed
- ▶ Dec 02 – Feb 03: Prototype Sector Line (PSL) & Mini Instrumentation Line deployed
- ▶ Mar 03: PSL and MIL connected to JB by submarine
- ▶ Jun 03: MIL & PSL recovered (diagnose: attenuation of the optical signal inside the mechanical electro optical cable of the lines)
- ▶ Mar 05: Line 0 & MIL deployed
- ▶ May 05 : Line 0 recovered (still attenuation of the optical signal inside the mechanical electro-optical cable of the lines due to a different reason)
- ▶ Oct 05: MIL still operating
- ▶ Oct 05: integration of Line 1 started with attenuation problem fixed
- ▶ Jan 06-Dec 07 : installation of the 12 lines



# The NEMO Project

## ► R&D phase (1999-2002)

- Site selection and characterization
  - Several sites close to the Italian coasts have been studied.
- R&D Activities
  - Development of dedicated ASICS for the underwater front-end electronics
  - Development of large area direction-sensitive optical modules
- Feasibility Studies
  - All detector critical components and the deployment procedures have been examined
  - A preliminary project for the km<sup>3</sup> detector has been developed

## ► Phase-1 and prototyping (2002-2006)

- Realization and deployment of a prototype including all critical components

## ► Phase-2 (2006-...)

- Realization of an underwater infrastructure at -3500 m

# NEMO R&D Activity

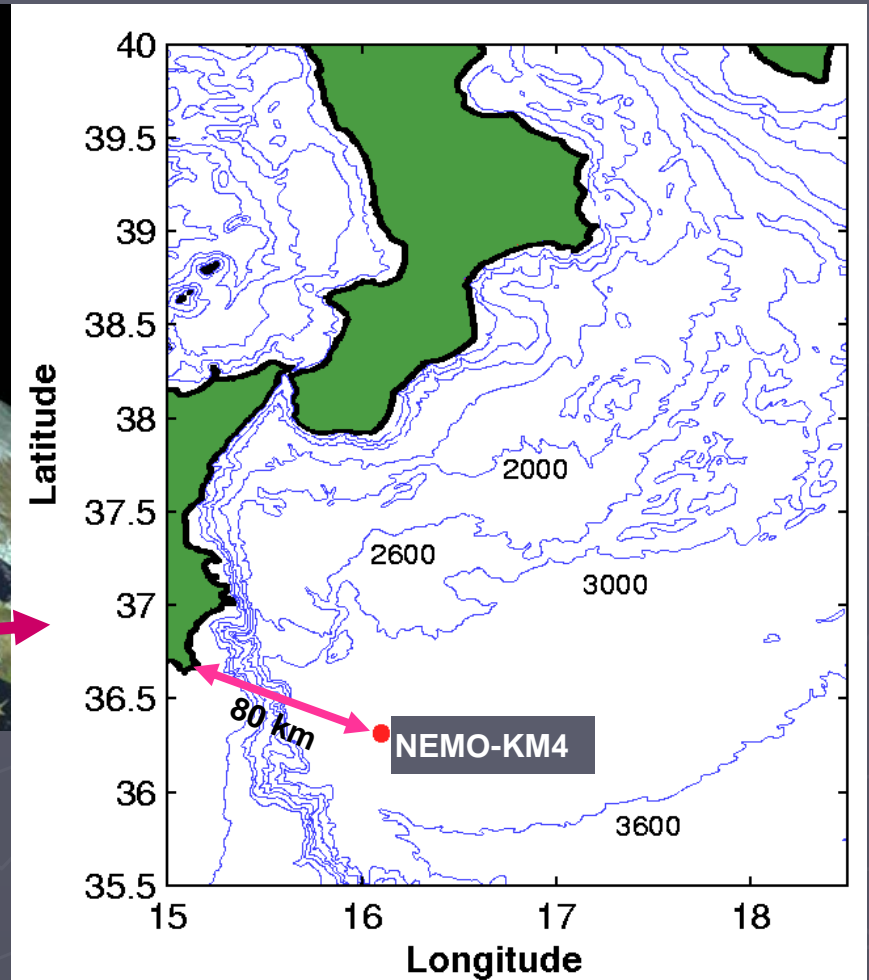
## ► Site location

- More than 25 sea campaigns since 1998



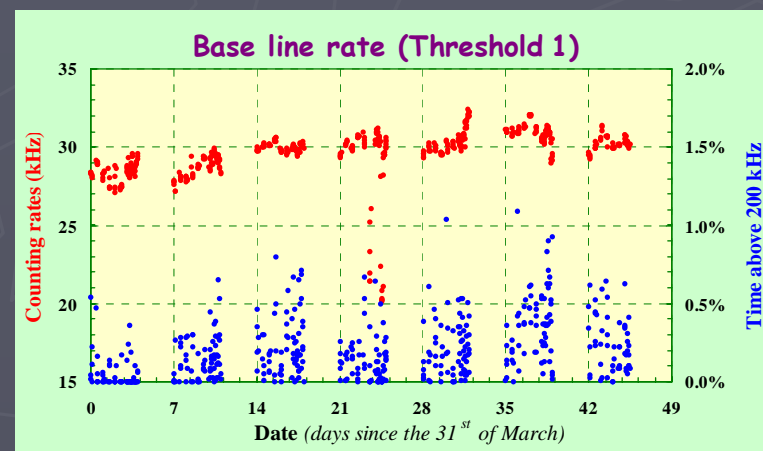
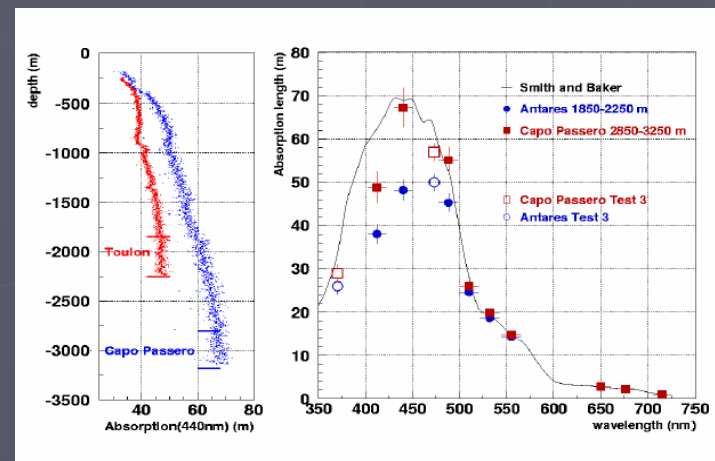
## ► NEMO-KM4

- 80 km off-shore
- 3500 m depth



# NEMO-KM4 Properties

- ▶ High water transparency
  - Data compatible with pure salt water properties
  - No seasonal variations
- ▶ Reduced background, mostly from  $^{40}\text{K}$  decay
  - 10" PMT thres. 0.5 p.e. noise rate  $\sim 30$  kHz
  - Bioluminescence almost absent
- ▶ Geologically stable

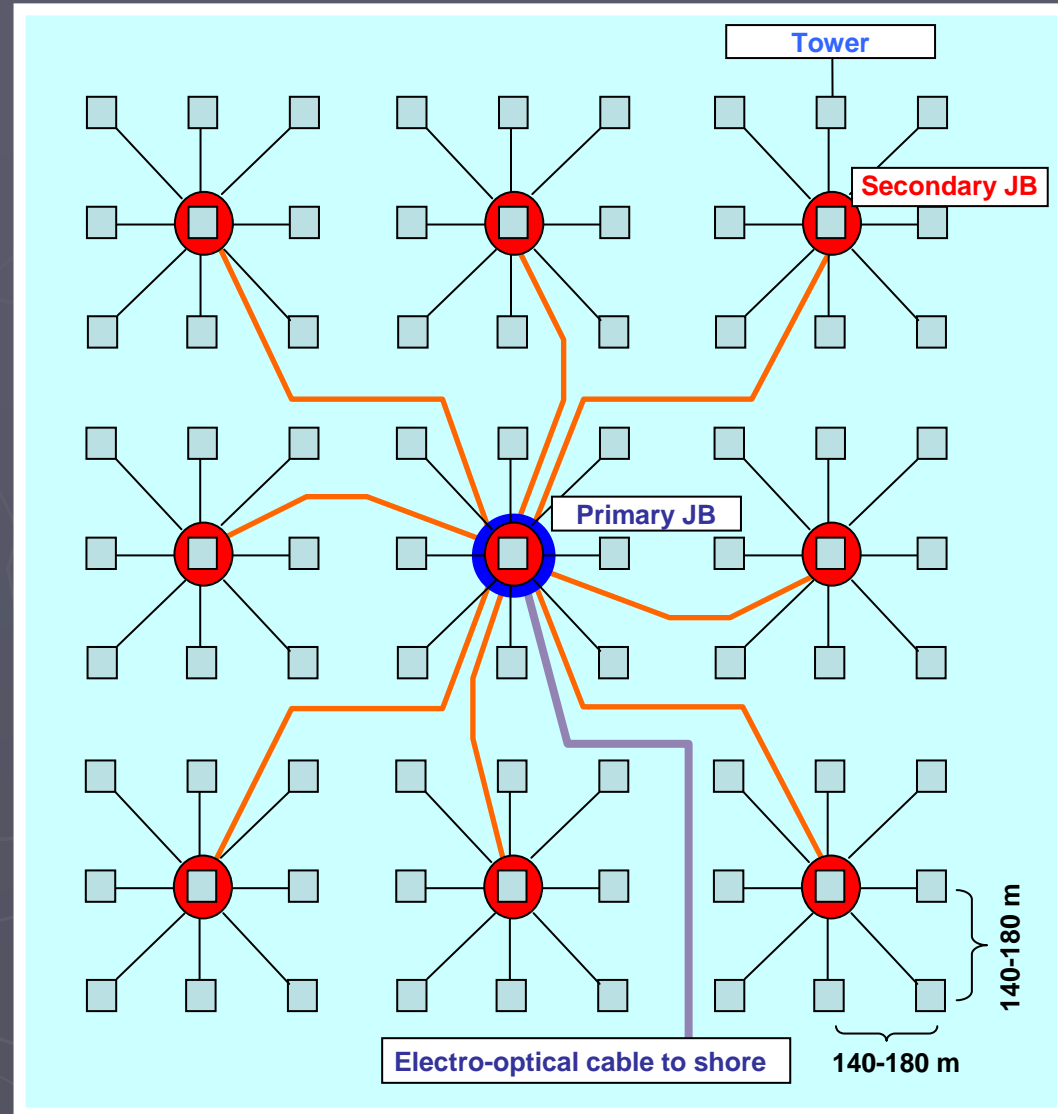


-3000 m

# NEMO km<sup>3</sup> Conceptual Design

## ► Proposed lay-out

- 10 junction boxes
- 81 towers
- 5832 PMTs



# Expected Performance

- Simulations show excellent angular resolution and sensitivity

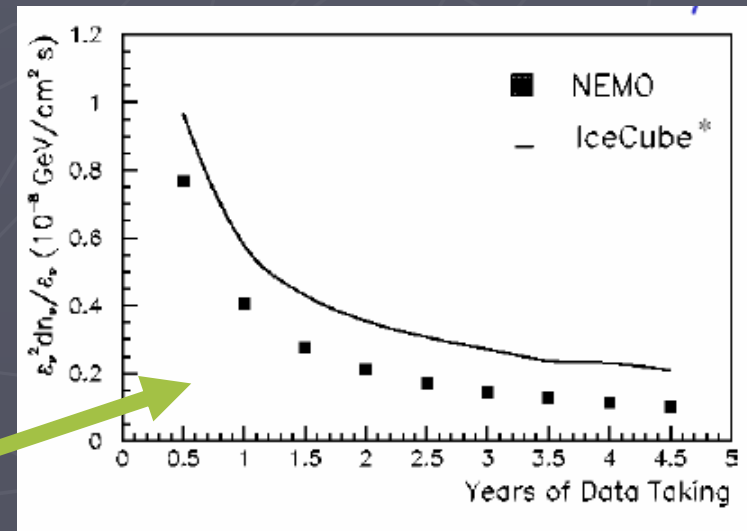
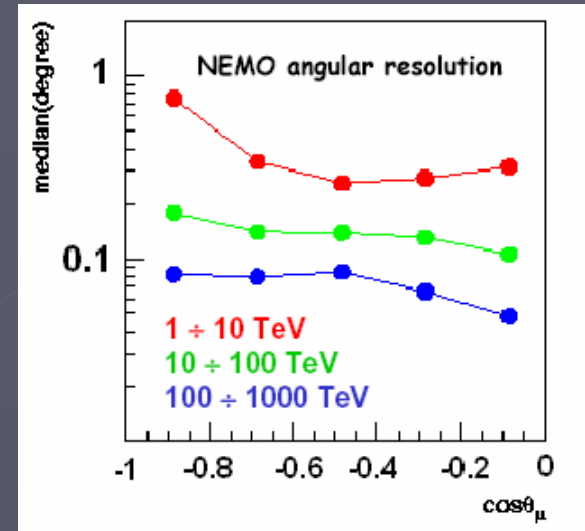
- NEMO

- 81 towers 140 m spaced
- 5832 PMTs

- ICECUBE

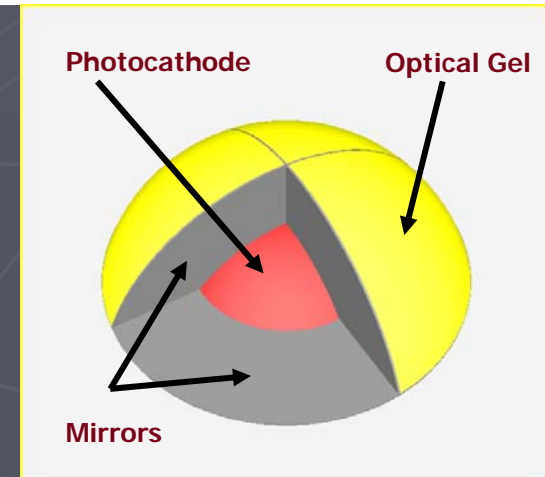
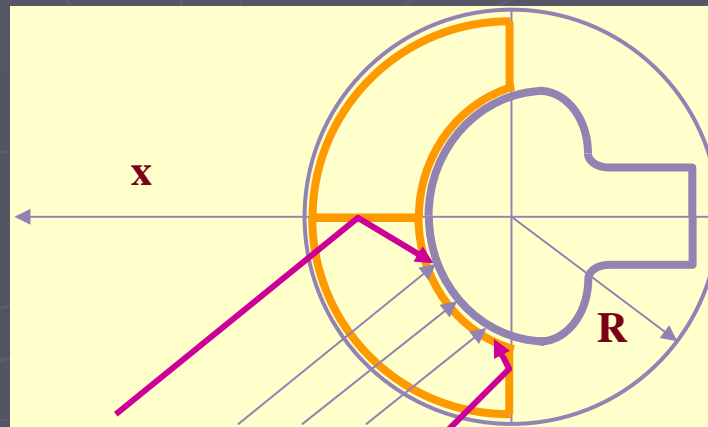
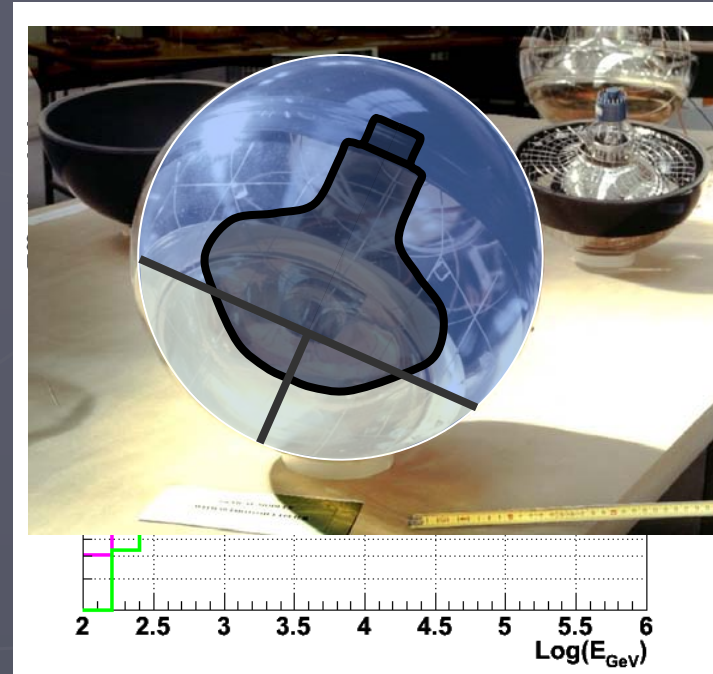
- 80 strings 125 m spaced
- 4800 PMTs

Sensitivity to a  $E^{-2}$  neutrino spectrum from a pointlike source



# Position-sensitive OMs

- ▶ Cherenkov light is emitted at fixed angle
  - Knowledge of direction of incoming light reduces background and uncertainties
- ▶ INFN Genova and MSU are collaborating to realize a prototype
  - 4 anode PMT
  - Mirror system





# The NEMO Phase-1 Activity

Shore station in the port of Catania



The Catania test site:

- 25 km off the coast of Catania
- 2000 m depth
- already equipped with a double-termination cable

Geoseismic station SN-1 (INGV)

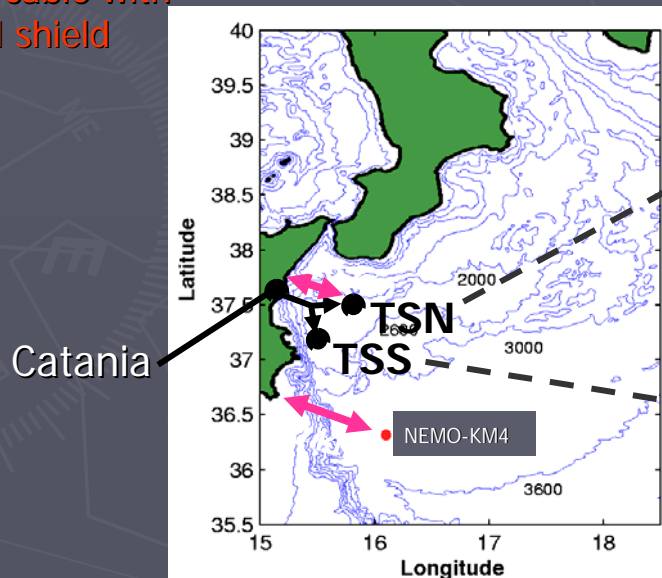


2.5 km e.o. cable with double steel shield

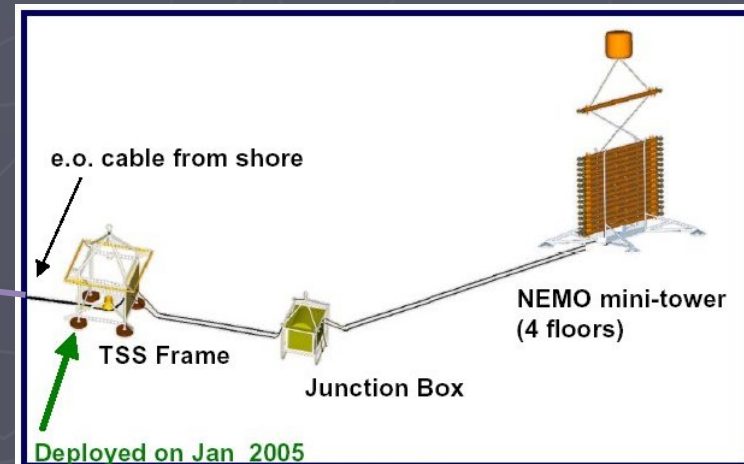
21 km e.o. cable with single steel shield

Branching Unit

5 km e.o. cable



5 km e.o. cable

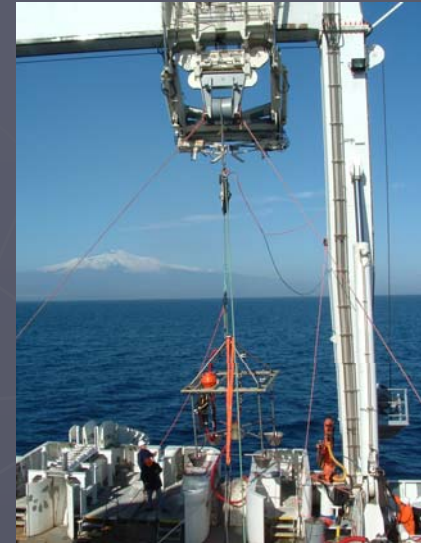
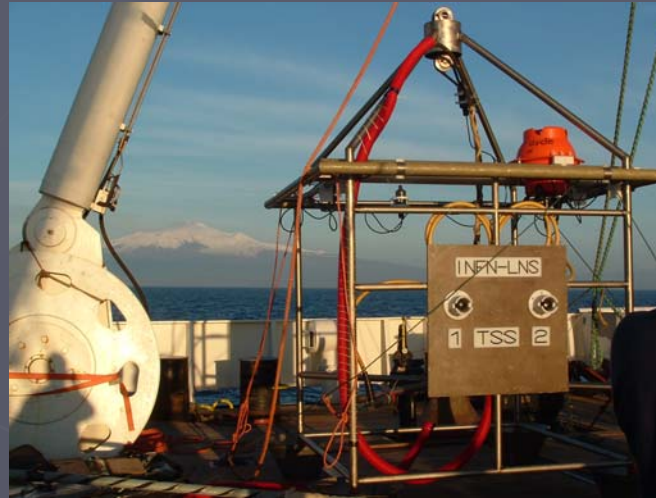


NEMO Phase 1



# The NEMO Phase-1 Activity

- Installation of the cable termination frames with electro-optical connectors

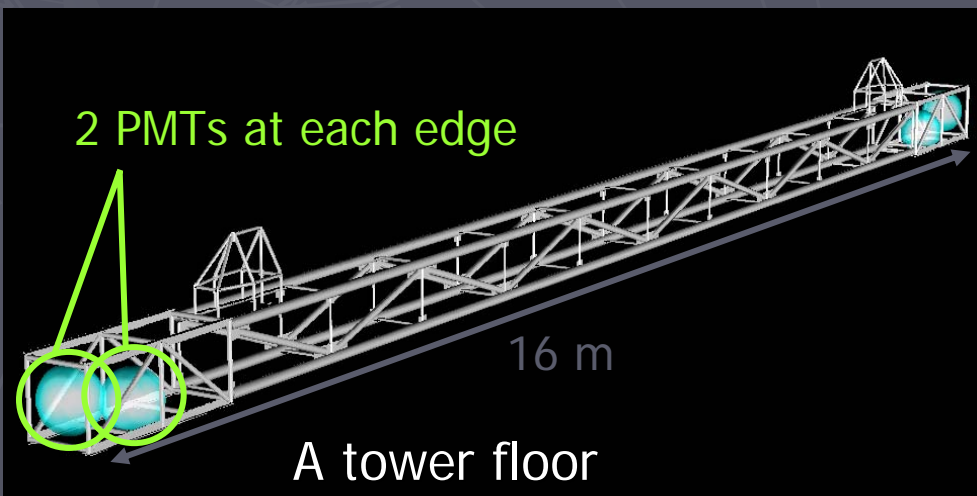


- Deployment and connection of
  - Acoustic detection station
  - INGV environmental observatory
- Fully operational since January 2005

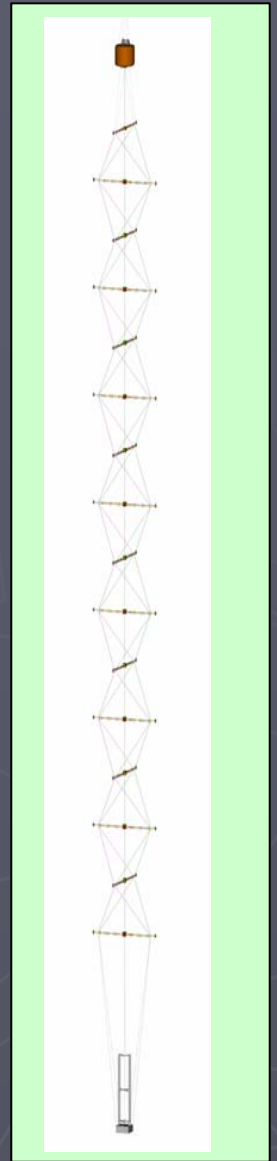
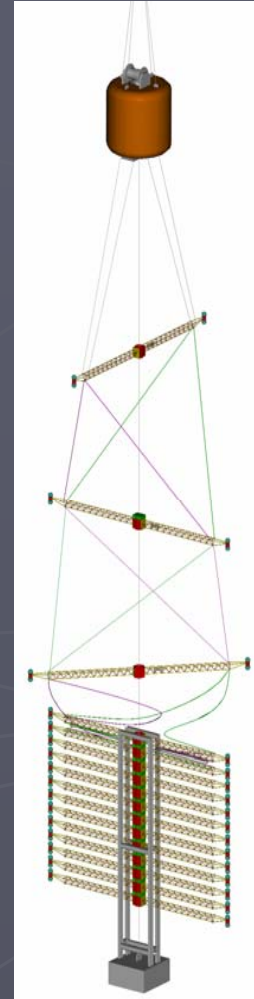
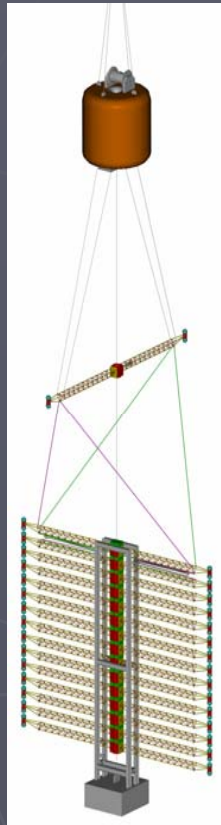
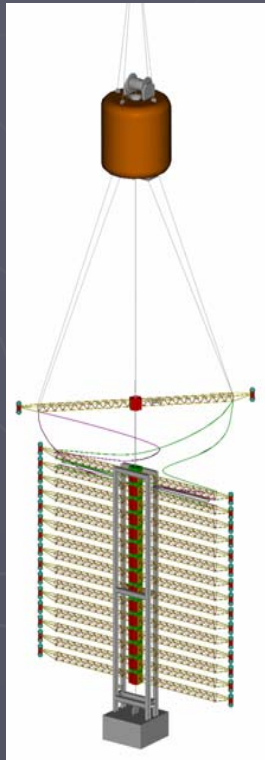
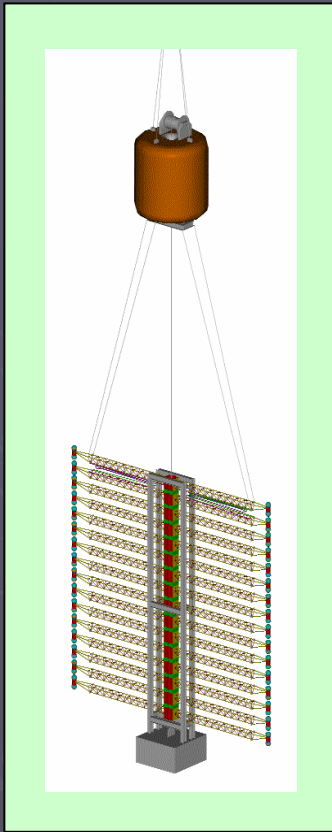


# The NEMO Tower

- ▶ Semi-rigid structure provides “easy” assembly, transportation and deployment
  - A 1:5 4-floor prototype has been successfully deployed and recovered in Spring 2004



# The Tower Deployment

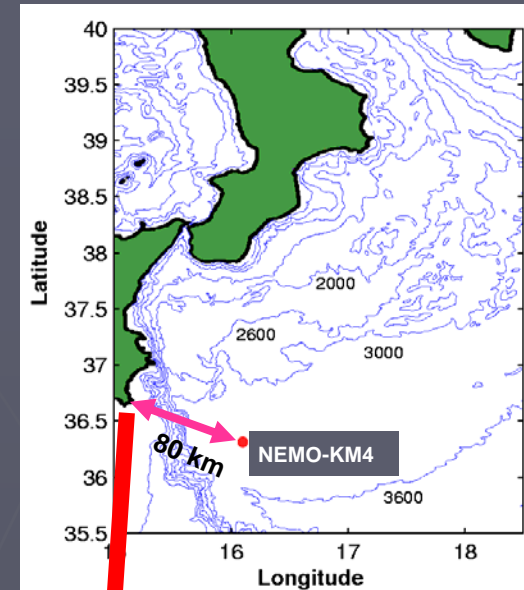




# The NEMO Phase-2

## ► Infrastructures for an -3500 m underwater site

- Electro-optical cable (40 kW)
  - Purchase under way
- Shore station in PortoPalo di Capo Passero
  - Building under renovation



to NEMO-KM4

# Conclusions and Perspectives

- ▶ The ANTARES Collaboration is realizing a small scale ( $0.1 \text{ km}^3$ ) neutrino detector in the Mediterranean Sea
  - Line 0 proved useful to solve problems related to data-transmission in high pressure environment
  - Line 1 is being integrated
  - It is expect to deploy the detector by the end of 2007

# Conclusions and Perspectives

- ▶ The NEMO Collaboration is working on a long-term R&D program toward a  $\text{km}^3$   $\nu$ -telescope in the Mediterranean Sea
  - An optimal candidate site has been found: NEMO-KM4
  - The NEMO Phase-1, aiming to validate the proposed technologies, is under way at the Catania Test Site
    - ▶ Since Jan. 2005 the geoseismic and acoustic stations are fully operative
    - ▶ The completion is planned in the first half of 2006
  - The NEMO Phase-2, aiming to realize the deep sea station at NEMO-KM4, is in progress
    - ▶ The purchase of the electro-optical cable is in progress
    - ▶ The set-up of the onshore station is in progress
    - ▶ The deployment of a full-size tower is foreseen in 2007

# Toward the $\nu$ -Telescope

- ▶ EU is funding the joint activity for an European-scale Design Study for a  $\text{km}^3$   $\nu$ -telescope in the Mediterranean Sea
  - KM3NeT: ANTARES-NEMO-NESTOR consortium
- ▶ 2° VLVnT (Very Large Volume  $\nu$ -Telescope) Workshop to be held in Catania (Italy) 8-11 Nov., 2005